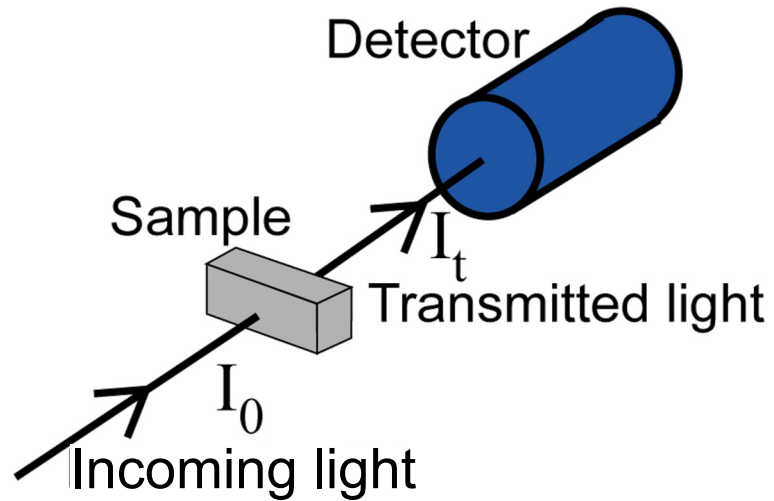


# ESC1: Circular Dichroism: best practice and data analysis

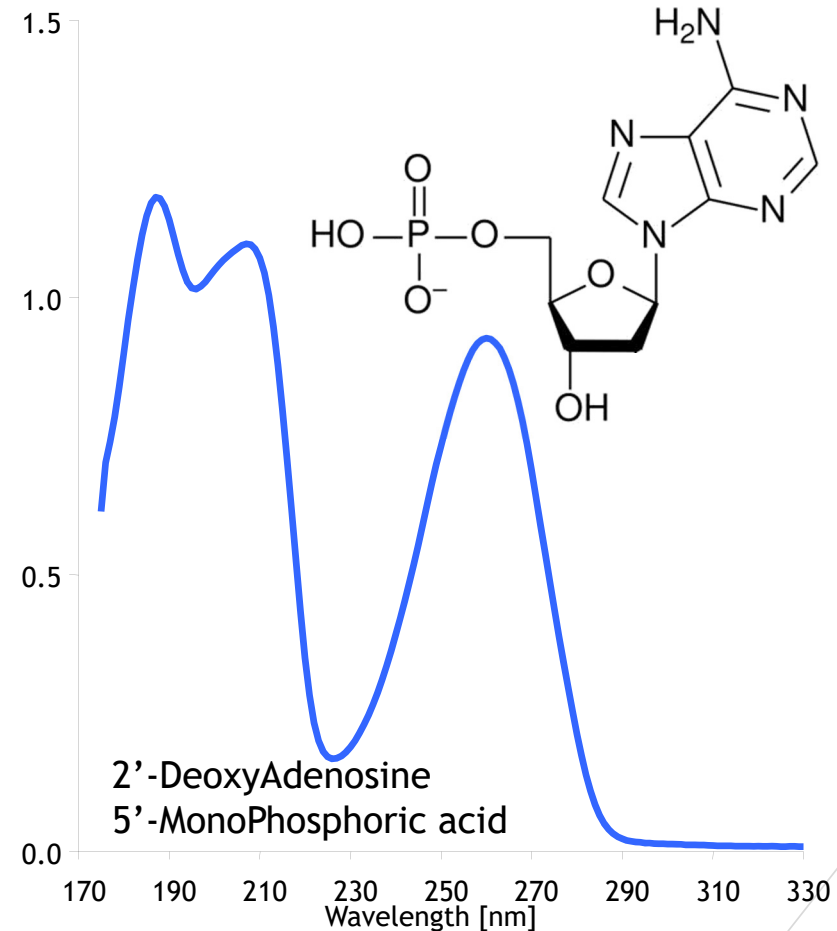
## Lecture 2: Instrumentation

# Absorption Spectroscopy

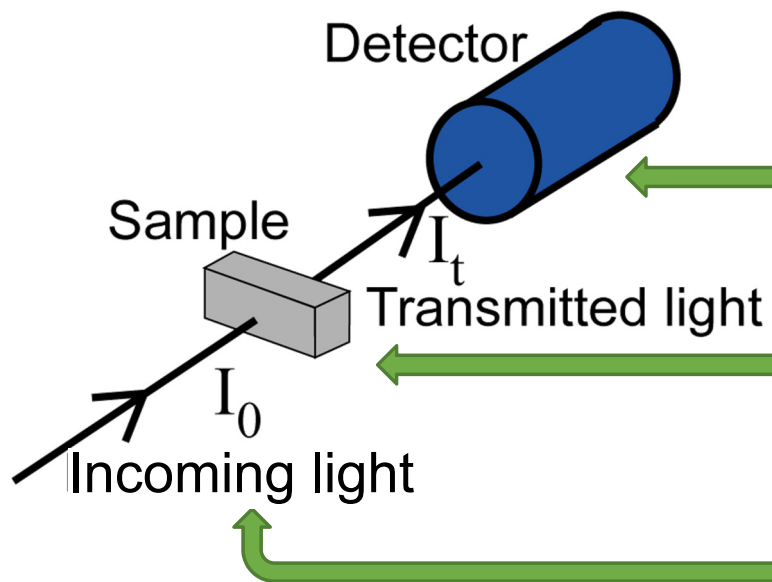


Absorbance:

$$A = \log(I_0 / I_t)$$



# UV-VIS Instrumentation



## Shopping list for making UV-VIS spectrometer:

- Detector
- Sample holder
- Light source with monochromatic light

*A monochromatic source means a single wavelength of light*

➡ To make such a source we need a ***lamp*** and ***optics***

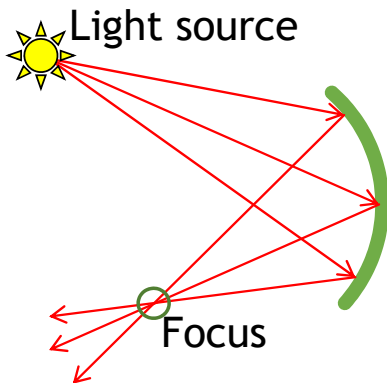
# UV-VIS Instrumentation

## Mirrors

Plane mirrors

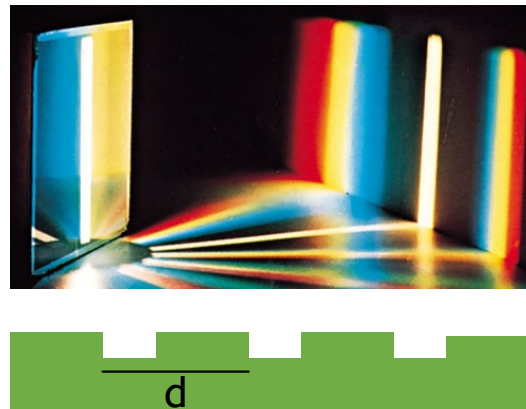


Curved mirrors

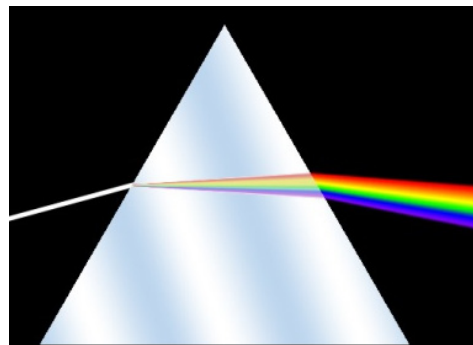


## Dispersing elements

Diffraction grating



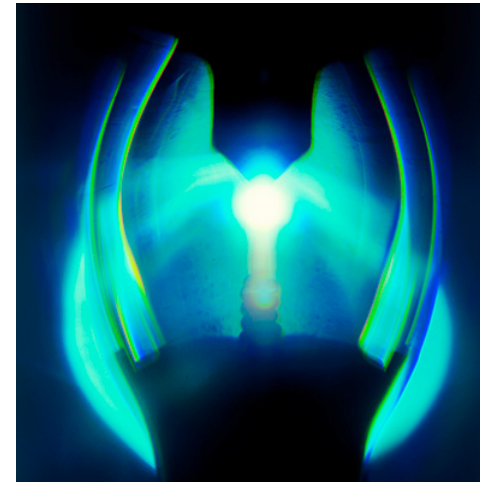
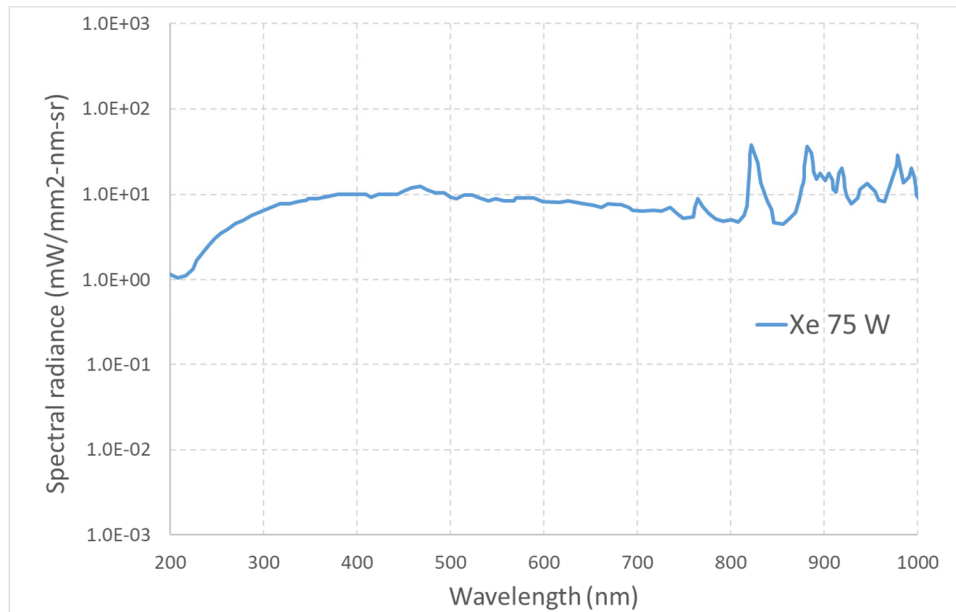
Prism





# UV-VIS Instrumentation: Light source

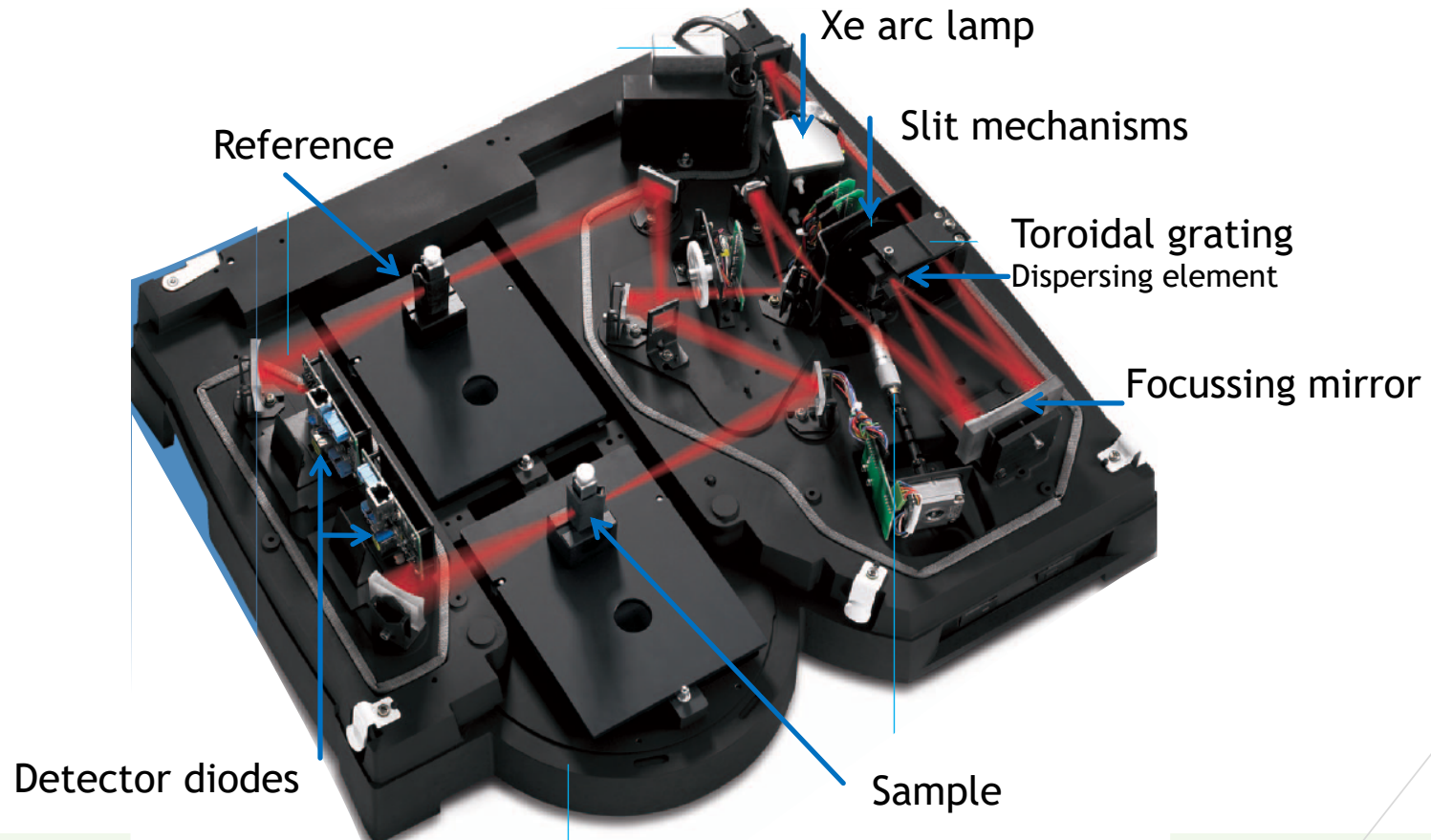
## Xenon arc lamp



*Quite intense UV lamp*

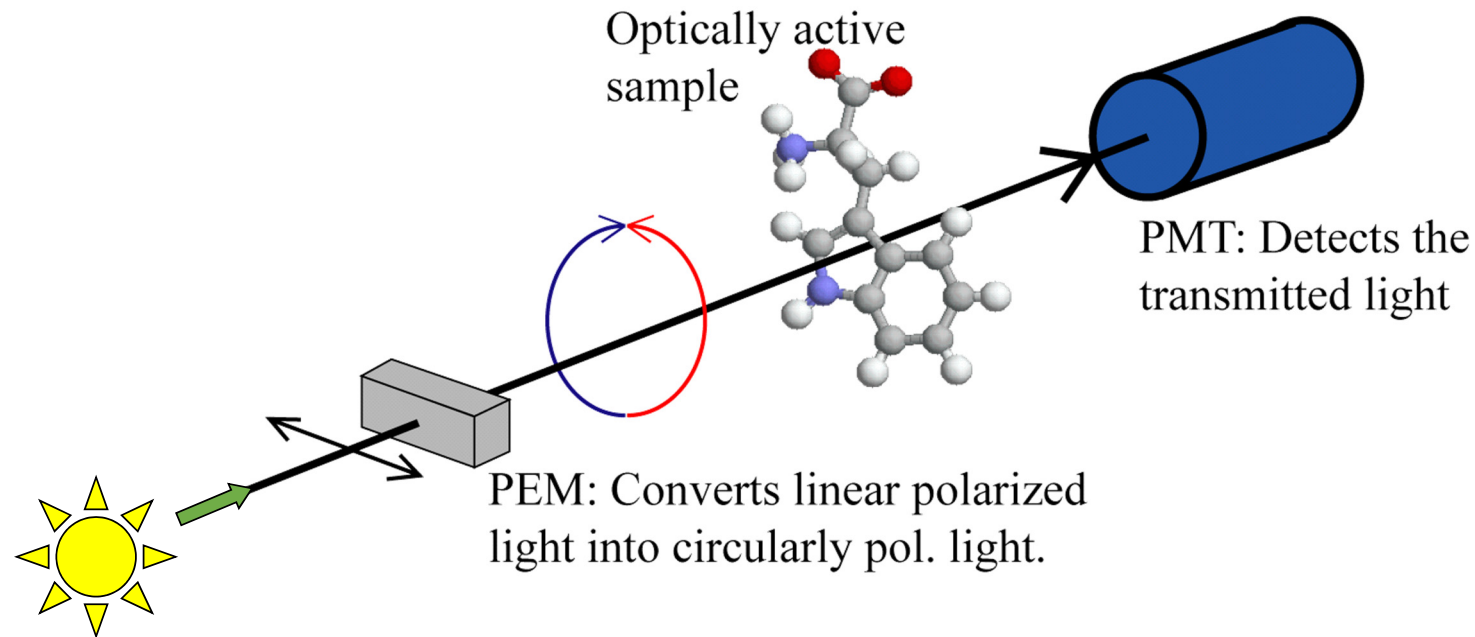
# UV-VIS Instrumentation

## Conventional UV-VIS instruments



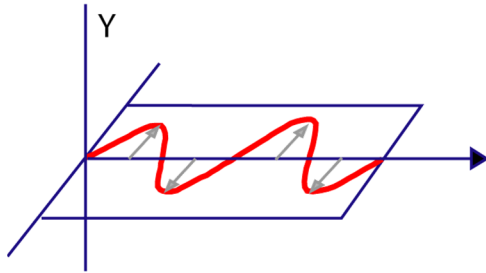
# Circular Dichroism

More optical components needed

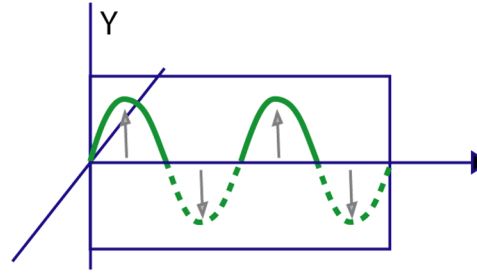


The CD signal:  $\Delta A = A_L - A_R$

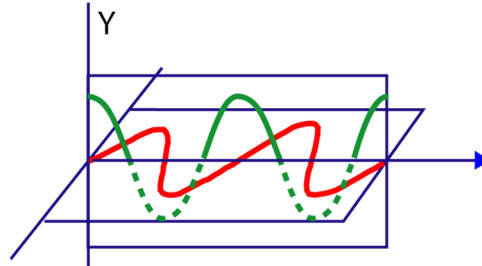
# Polarized light



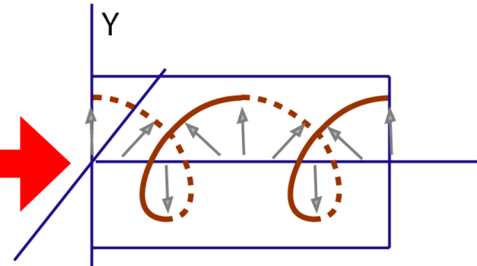
Horizontally pol.



Vertically pol.



Sum of two  
plane pol.

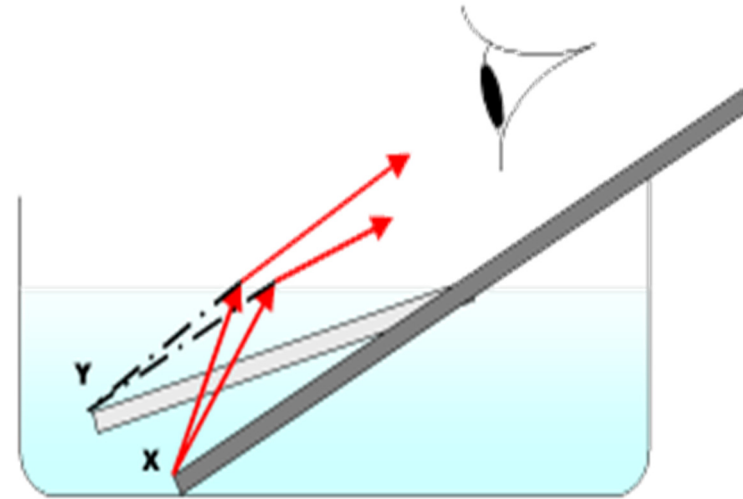


Circularly pol.

How do we change the polarisation of light ?

# Refraction

Refraction: Light is refracted when travelling from one medium to the next



This is (partly) why it's difficult to catch a fish in water

# Refraction

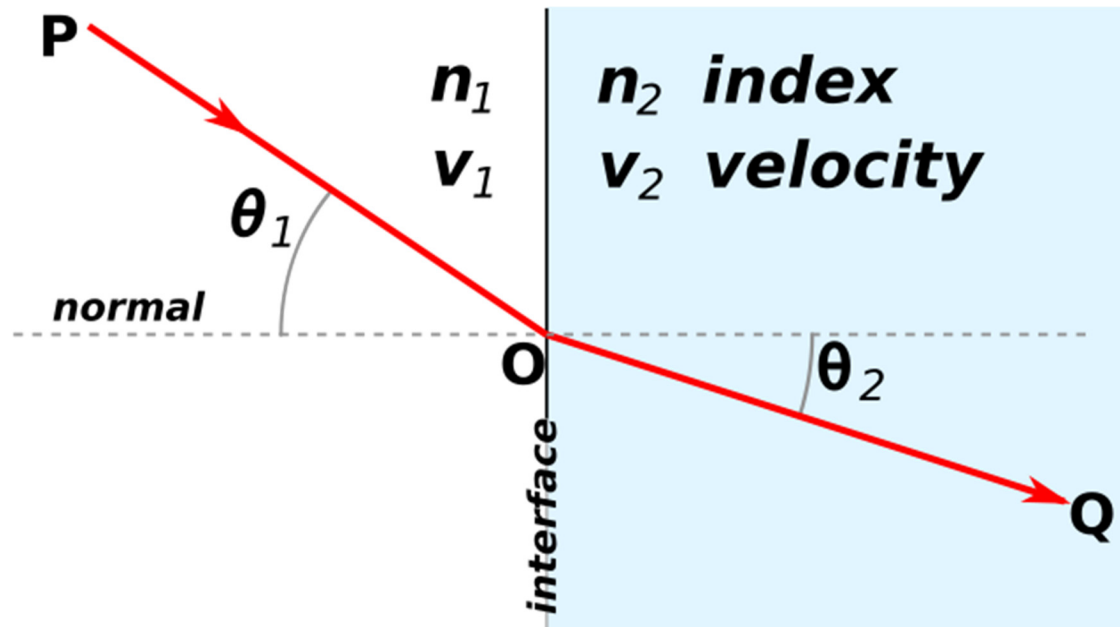
Snell's law

$$\frac{\sin \theta_1}{\sin \theta_2} = \frac{n_2}{n_1}$$

Index of refraction

$$n = \frac{c}{v_p}$$

The speed of light in a material depends on the properties of the material





# Birefringence

Birefringence, or double refraction



Icelandic Spar / Calcite /  $\text{CaCO}_3$

Used by the Vikings for navigation (sunstone)

- To tell the direction of the sun on a cloudy day



# What is Birefringence?

Indexes of refraction depend on polarization

The birefringent effect (using calcite) was first described by scientist Rasmus Bartholin in 1669

If a crystal has two different Indexes of refraction:  $n_e$  and  $n_o$

**Birefringence**

$$\Delta n = n_e - n_o$$

Calcite:  $\Delta n(590\text{nm}) = 1.486 - 1.64 = - 0.154$





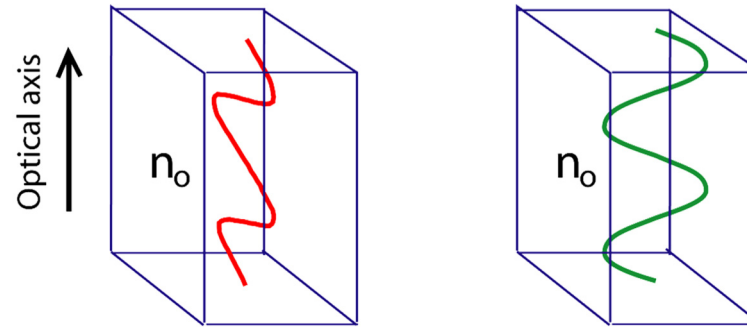
# What is Birefringence

## What is meant by two indexes of refraction?

The material (crystal) has a direction: the optical axis

Light travelling along the optical axis:

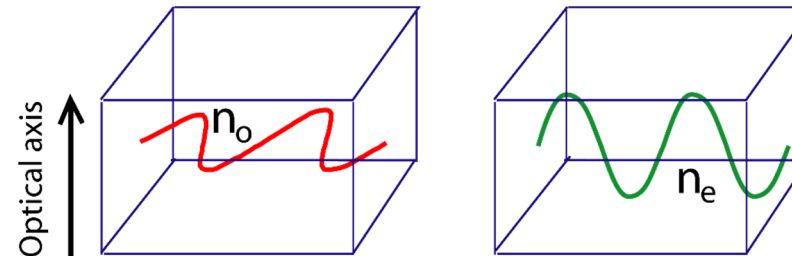
$n$  independent of polarization



Light travelling perpendicular to the optical axis:  
 $n$  depends on polarization

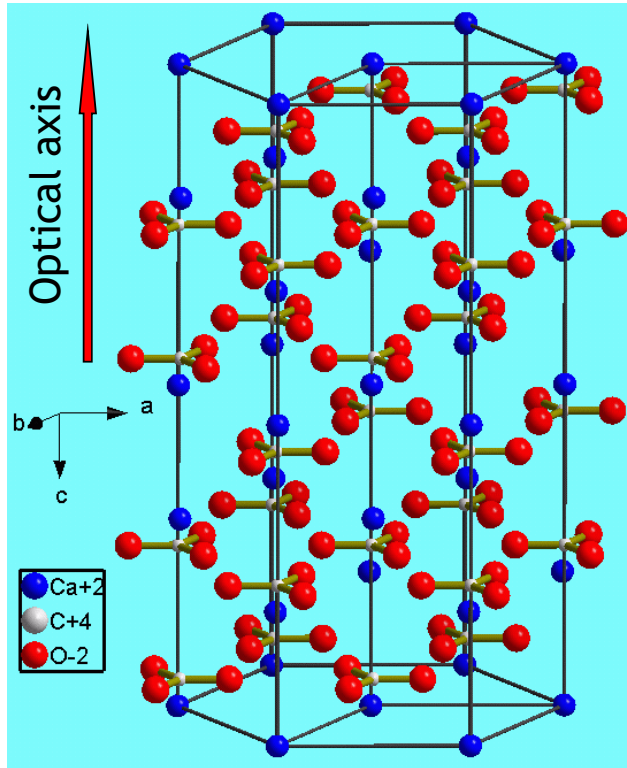
Pol.  $\perp$  optical axis:  $n_o$

Pol.  $\parallel$  optical axis:  $n_e$

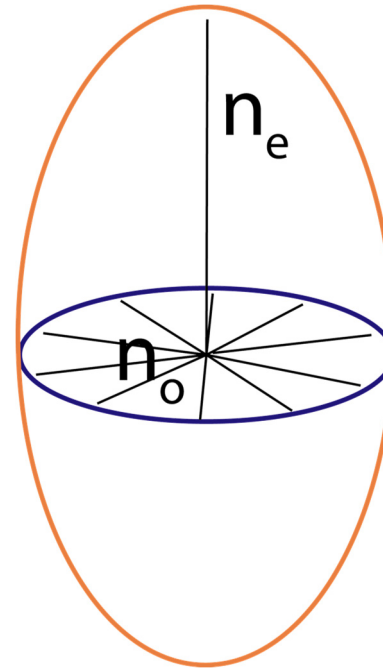


# Origin of Birefringence

## Crystal structure of Calcite



Calcite is an uniaxial crystal



Birefringence

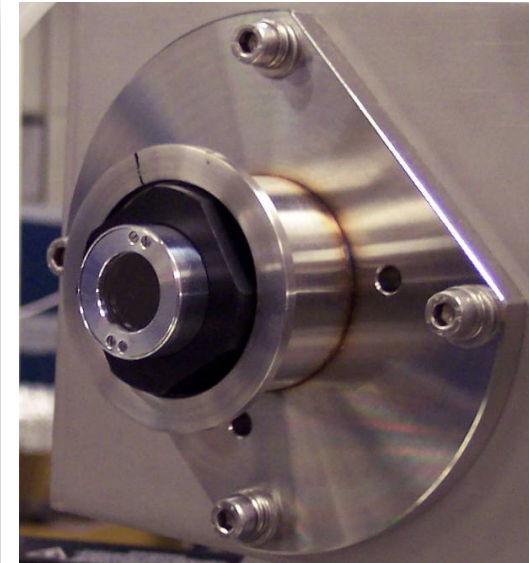
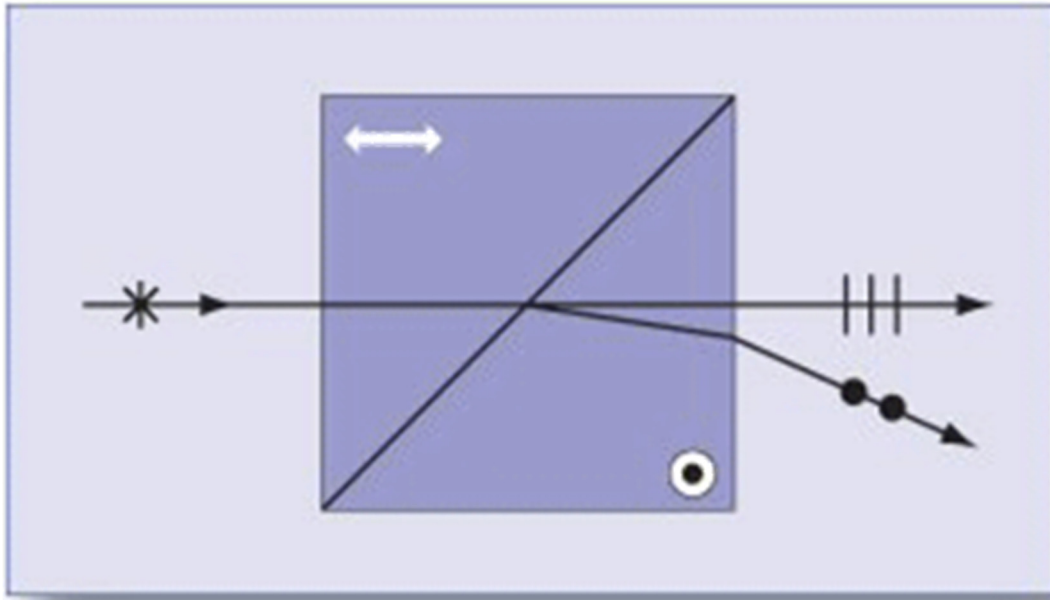
$$\Delta n = n_e - n_o$$

# The Rochon Polarizer

*“MgF<sub>2</sub> is slightly birefringent...”*


- Info on MgF<sub>2</sub> from Crystran Ltd., UK.

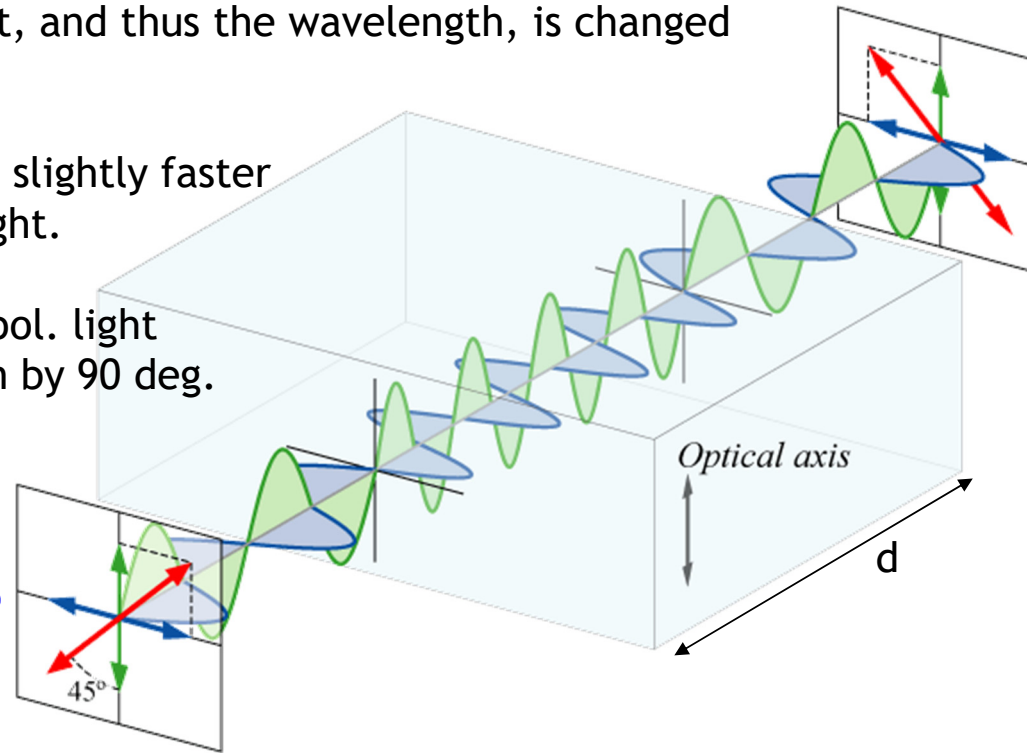
$$\text{MgF}_2: \Delta n(546\text{nm}) = 1.390 - 1.379 = 0.011 \quad (<< \Delta n_{\text{Calcite}})$$



# Retarders/Wave plates

- The speed of the light, and thus the wavelength, is changed inside the crystal
- The horz. pol. light is slightly faster than the vert. pol. light.
- The initially 45 deg. pol. light 'flips' the polarization by 90 deg.

  
 $\frac{1}{2}$  waveplate

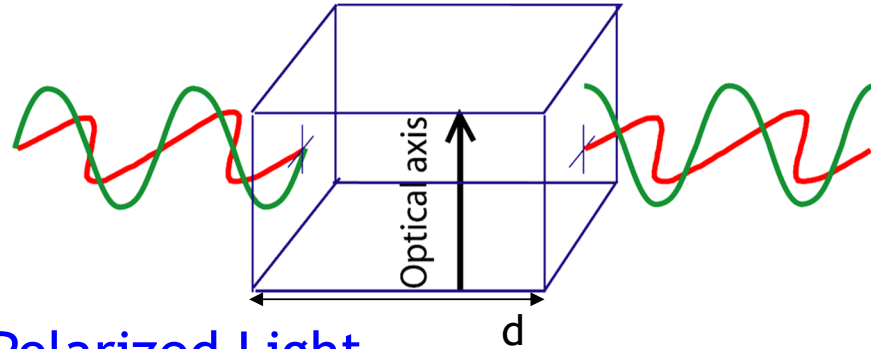


Retardation

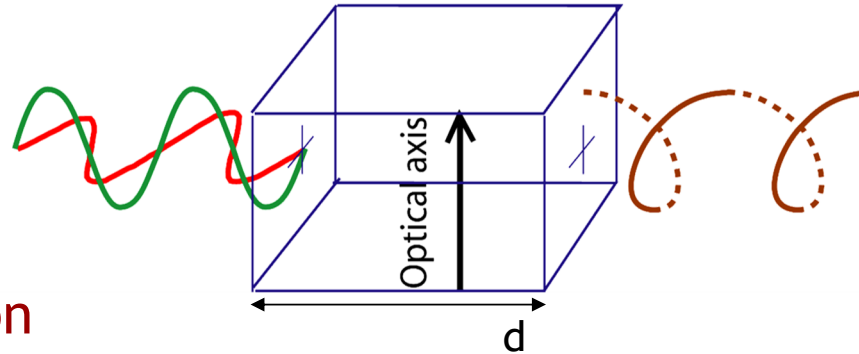
$$R = (n_e - n_o) d / \lambda$$

# Retarders/Wave plates

Half the thickness:   $\frac{1}{4}$  waveplate



Circularly Polarized Light

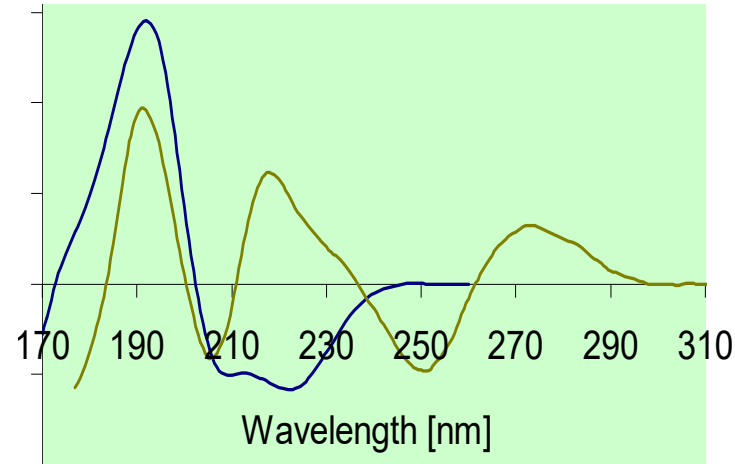


Retardation

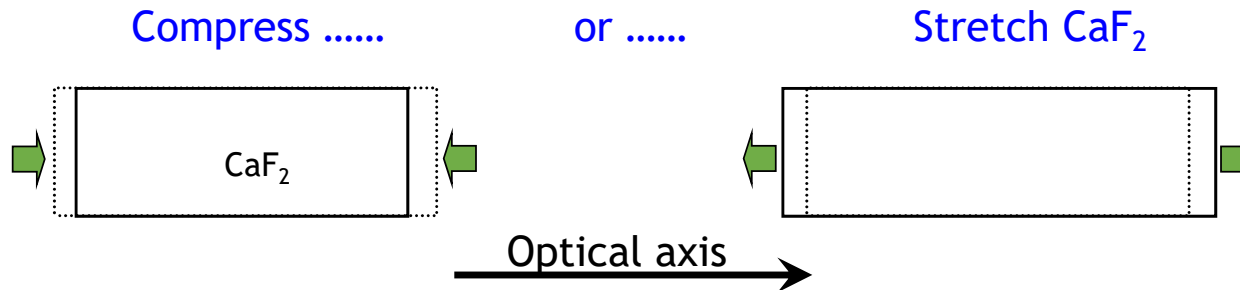
$$R = (n_e - n_o) d / \lambda = \frac{1}{4}$$

# Photo Elastic Modulators (PEM)

- We are scanning the wavelength
- We can't change  $d$  as we scan

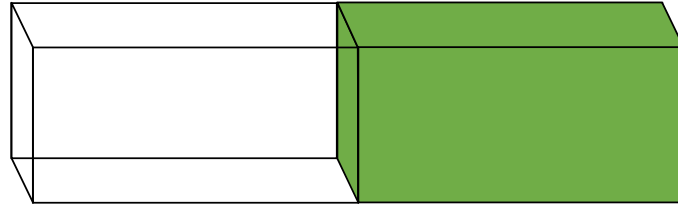


To keep  $R = 1/4$  (circularly pol)  
change  $n_e - n_o$  by:



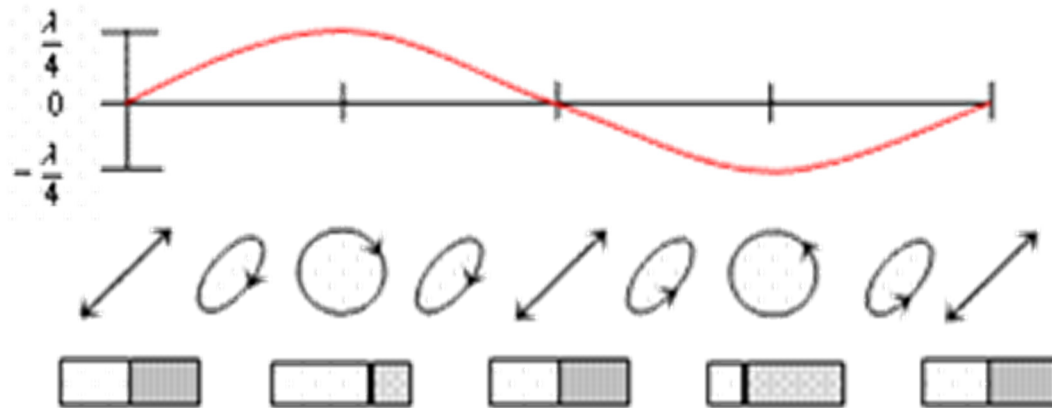
# Photo Elastic Modulators (PEM)

CaF<sub>2</sub> crystal



Piezo crystal.

Vibrates at ~50 kHz

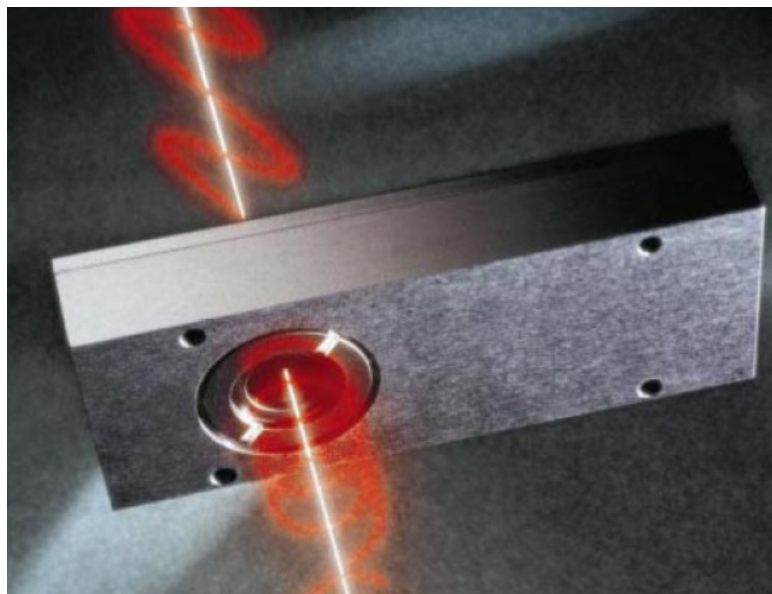


As the crystal vibrates the polarization is changed:

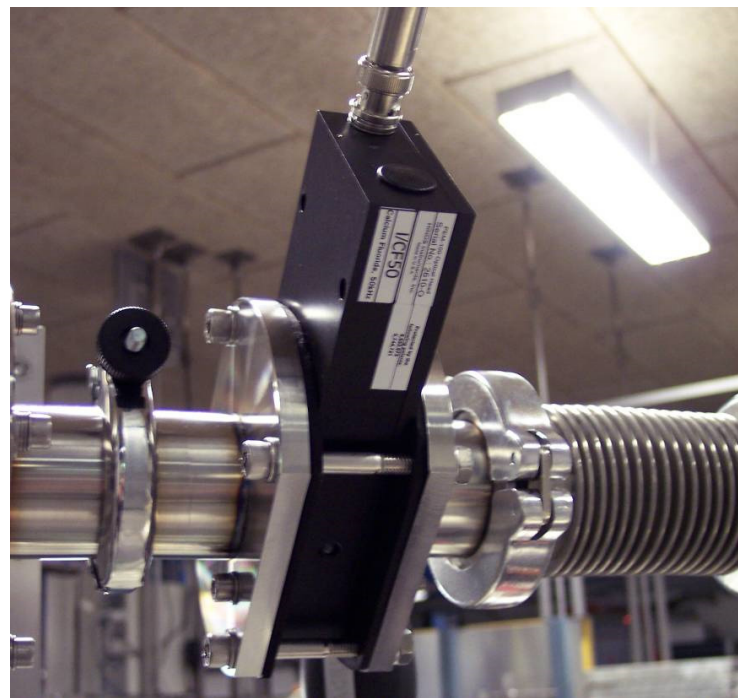
Lin.  $\longrightarrow$  Left Circ.  $\longrightarrow$  Lin.  $\longrightarrow$  Right Circ.  $\longrightarrow$  Lin.  $\longrightarrow$  ...



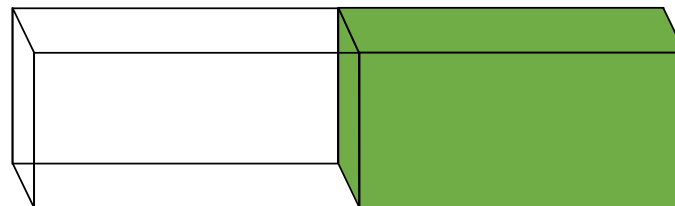
# Photo Elastic Modulators (PEM)



Linear pol. is converted to Circ. pol.



$\text{CaF}_2$  crystal



Piezo crystal.

Vibrates at ~50 kHz

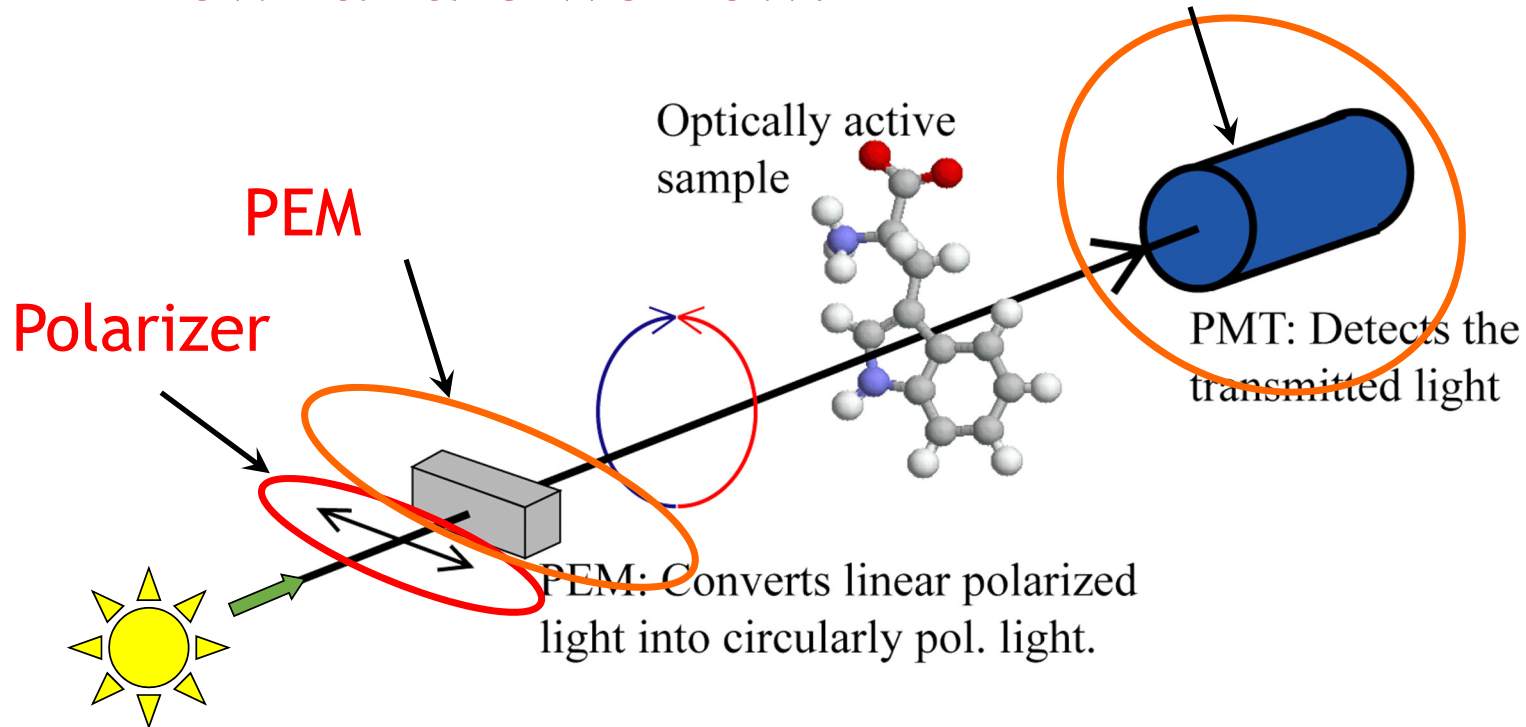




# Circular Dichroism

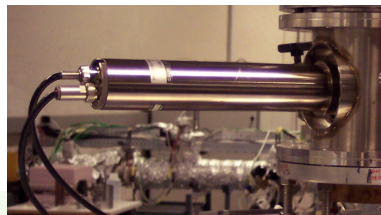
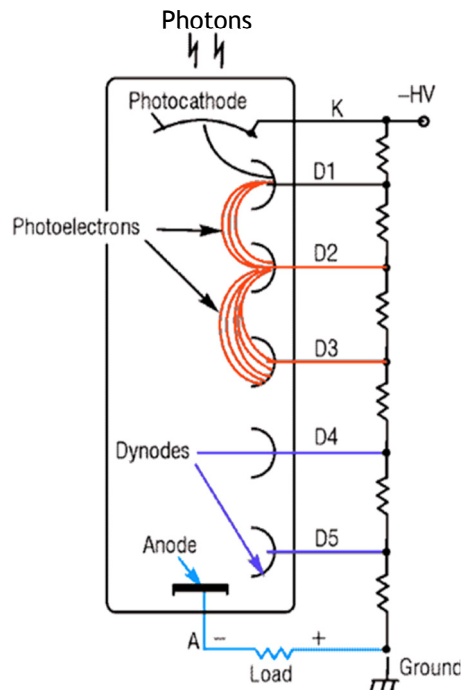
How far are we now?

Next: Detector



# CD/UV-VIS Instrumentation

## Photomultiplier Tube (PMT) Detector

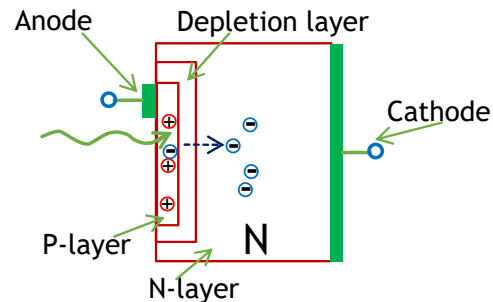
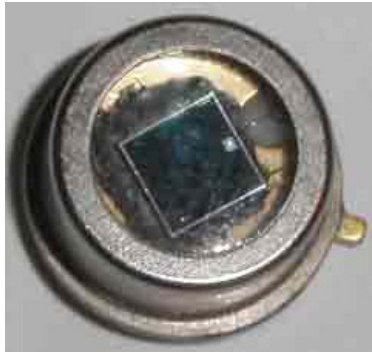


- A photon is converted to an electron on the cathode:  
 $Q_{\text{eff}}$  (quantum efficiency)
- The cathode can be optimized to certain wavelength ranges  
UV, VIS or even Solar Blind
- A high voltage (HV) drop is applied along the dynodes
- The photon electron is amplified to many electrons:  
 $\text{Gain(HV)}$  (strongly depends on HV)  
 $\text{Gain(HV)} \approx b \cdot (\text{HV})^a$
- The electrons are collected on the Anode:  
Either pulses or a **current** is detected  
Typically:  $\text{Gain(HV)} \approx 10^3 - 10^7$  and  $Q_{\text{eff}} < 1$

$$I_{\text{Anode}}(\lambda, \text{HV}) = \text{Flux [ph/sec]} \cdot e \cdot Q_{\text{eff}}(\lambda) \cdot \text{Gain(HV)}$$

# CD/UV-VIS Instrumentation

## Photodiode



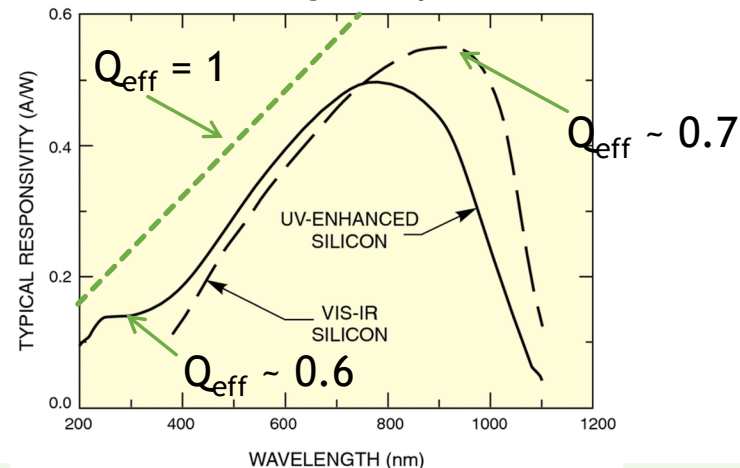
- The diode converts photons to a current
- $Q_{\text{eff}}$  can be as high as 80%
- Spectral range depends on material  
E.g. Si 190 - 1100 nm, Ge 400 - 1700 nm

### Advantage:

Very rugged, high  $Q_{\text{eff}}$  (in the VIS/IR)

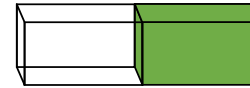
### Disadvantage:

No gain, although avalanche diodes can have a gain up to 1000



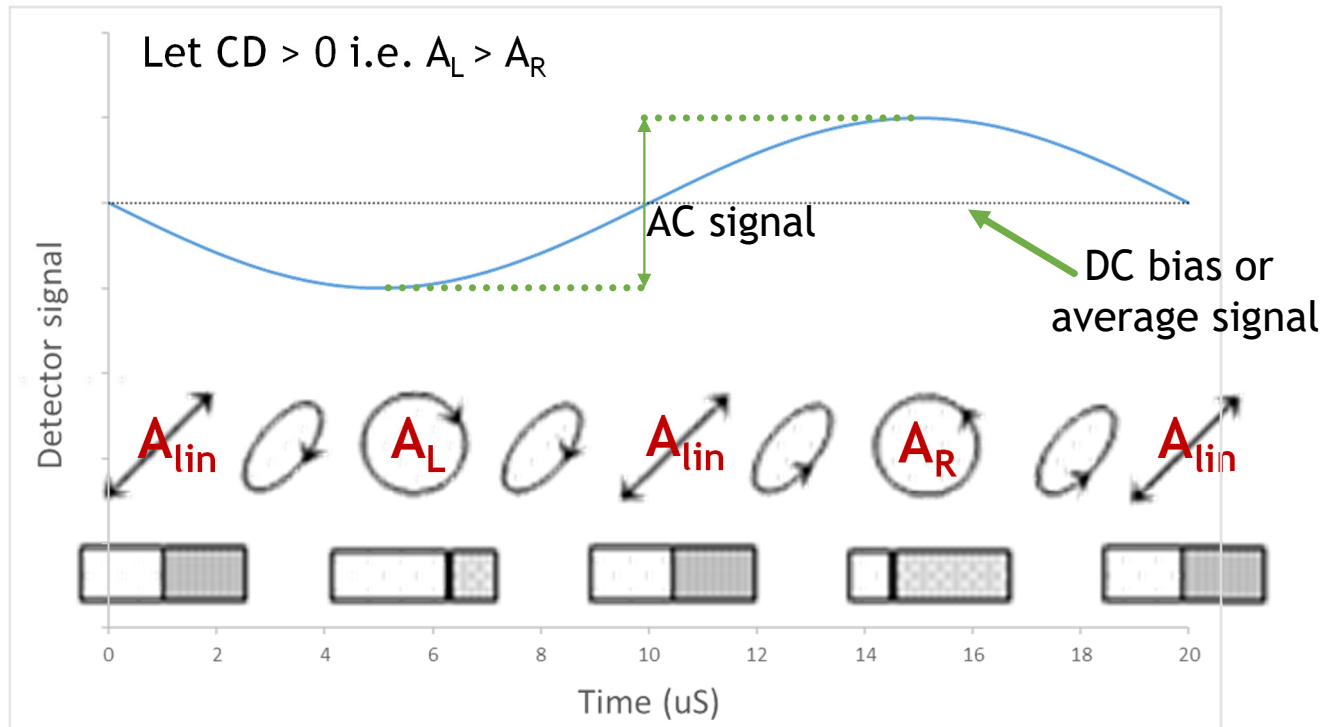
# The signal from the detector

The PEM oscillates and changes the polarization

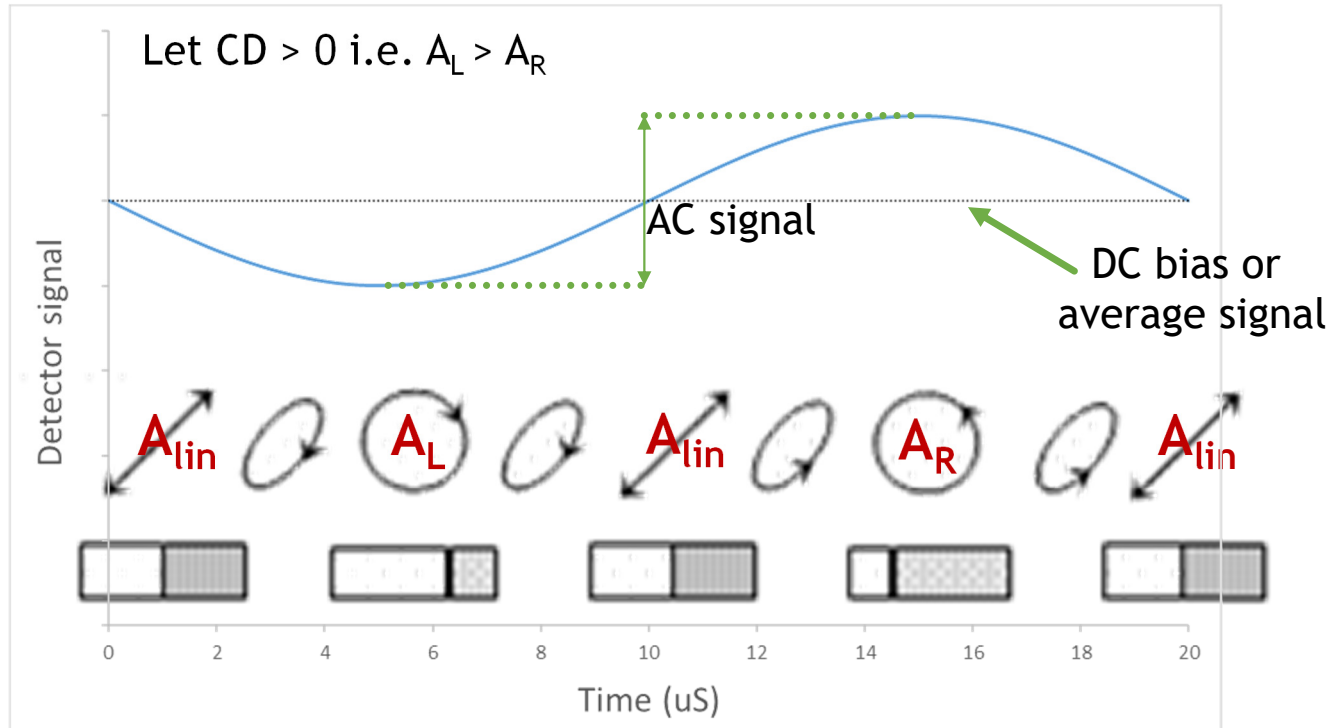


Vibrates at ~50 kHz

- Polarization L  $\rightarrow$  absorbance  $A_L$
- Polarization R  $\rightarrow$  absorbance  $A_R$
- Polarization is linear  $\rightarrow A_{lin} = \frac{1}{2}(A_L + A_R)$



# The signal from the detector



The CD instrument measures the AC (and DC)

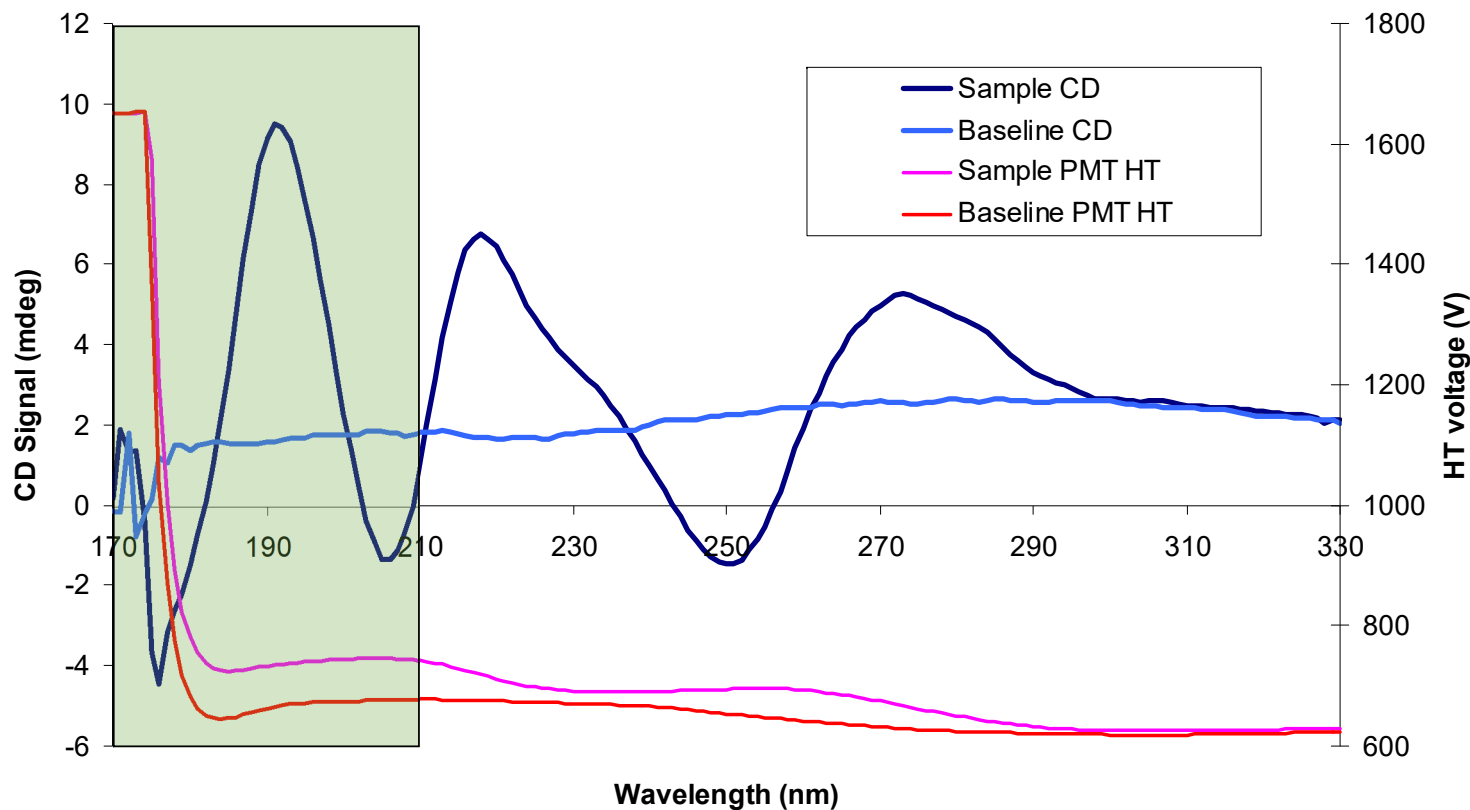
$$CD = constant \times \frac{AC}{DC}$$

*In most instruments the DC is kept constant by changing the detector high voltage (HT)*

$$CD = const \times AC$$

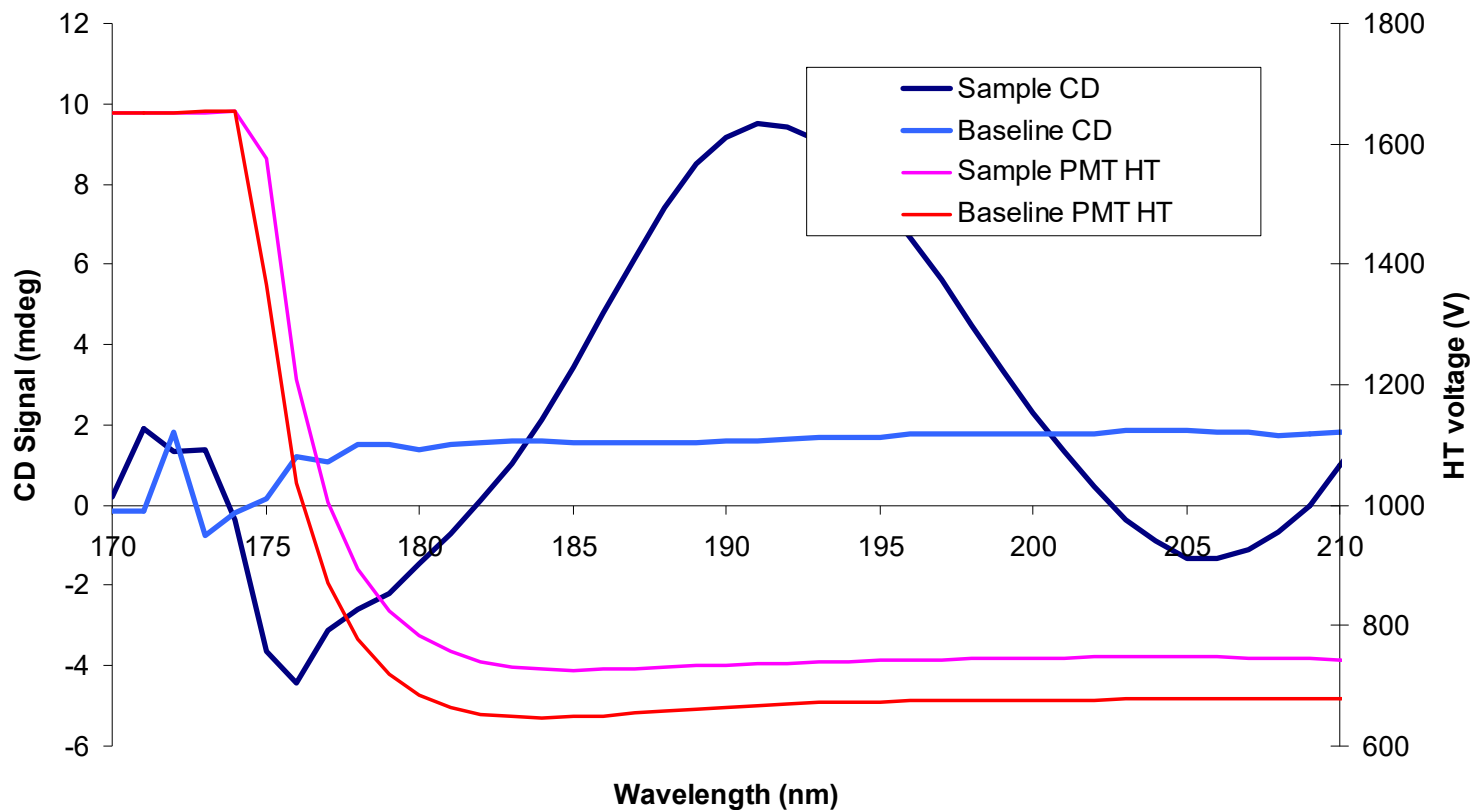
# What does the HT tell us?

High HT  $\longleftrightarrow$  Low photon flux reaching detector  $\longleftrightarrow$  High Absorbance



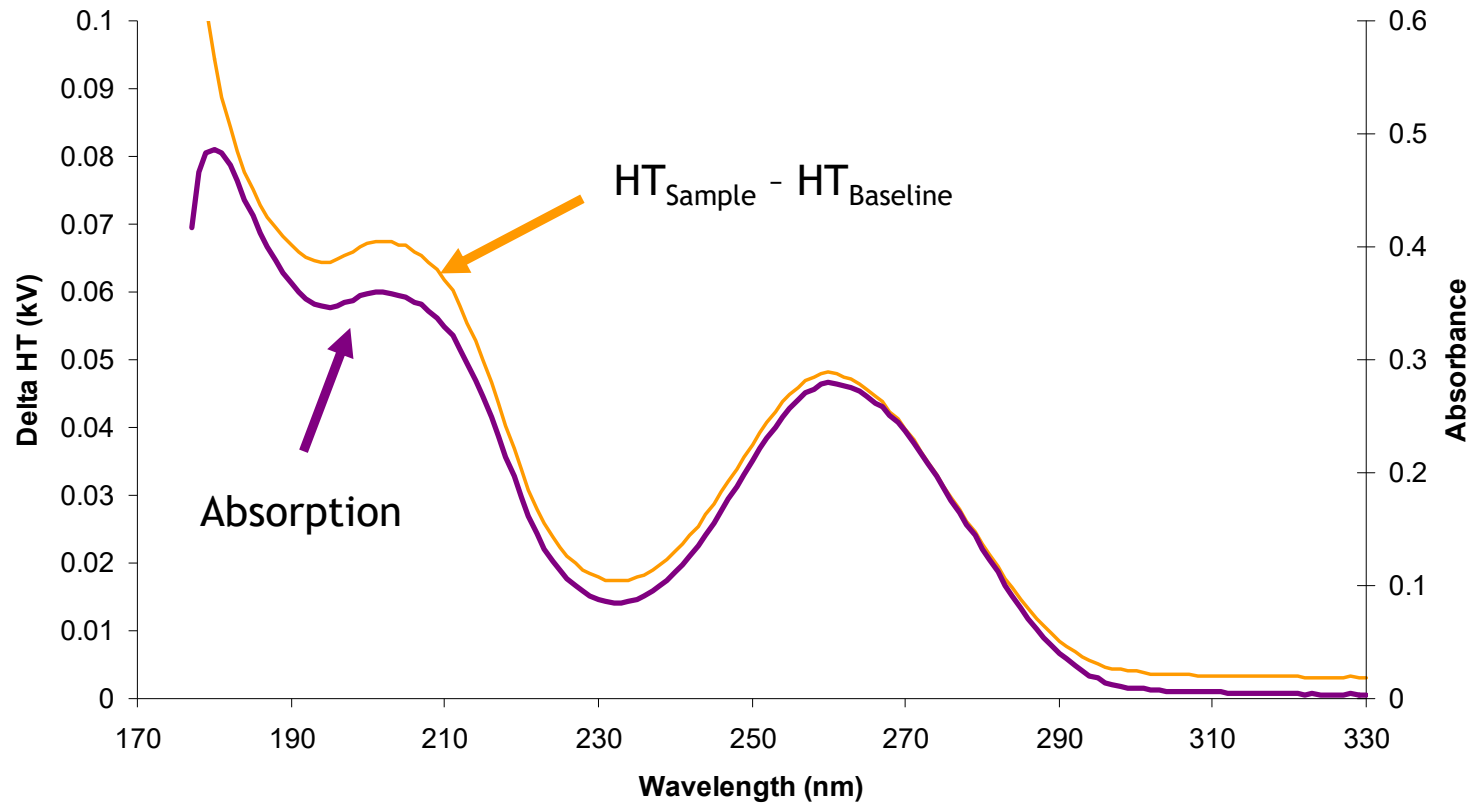
# What does the HT tell us?

High HT  $\longleftrightarrow$  Low photon flux reaching detector  $\longleftrightarrow$  High Absorbance



# What does the HT tell us?

Delta HT ( $HT_{\text{Sample}} - HT_{\text{Baseline}}$ ) is a 'pseudo' absorption





# What does the HT tell us?

It is possible to use  $HT_{\text{Sample}}$  and  $HT_{\text{Baseline}}$  to calculate the absorbance

- Record the  $HT_{\text{Sample}}(\lambda)$  and  $HT_{\text{Baseline}}(\lambda)$  together with the CD signal
- The average signal from the detector is constant (DC bias)
- Assume that for sample and baseline these are the same:
  - *Lamp output*
  - *Optics transmission*
- Assume the gain of the detector vs. HT is  $\text{Gain}(HT) = b \times HT^a$

$$\text{Detector signal}_{\text{sample}} = \text{Detector signal}_{\text{baseline}}$$

$$10^{-A_{\text{sample}}(\lambda)} \times \text{Gain}(HT_{\text{sample}}(\lambda)) = 10^{-A_{\text{baseline}}(\lambda)} \times \text{Gain}(HT_{\text{baseline}}(\lambda))$$

Use  $\text{Gain}(HT) = b \times HT^a$  :  $10^{A_{\text{sample}}(\lambda) - A_{\text{baseline}}(\lambda)} = \frac{b \times (HT_{\text{sample}}(\lambda))^a}{b \times (HT_{\text{baseline}}(\lambda))^a}$

$$A(\lambda) = a \times \log\left(\frac{HT_{\text{sample}}(\lambda)}{HT_{\text{baseline}}(\lambda)}\right)$$

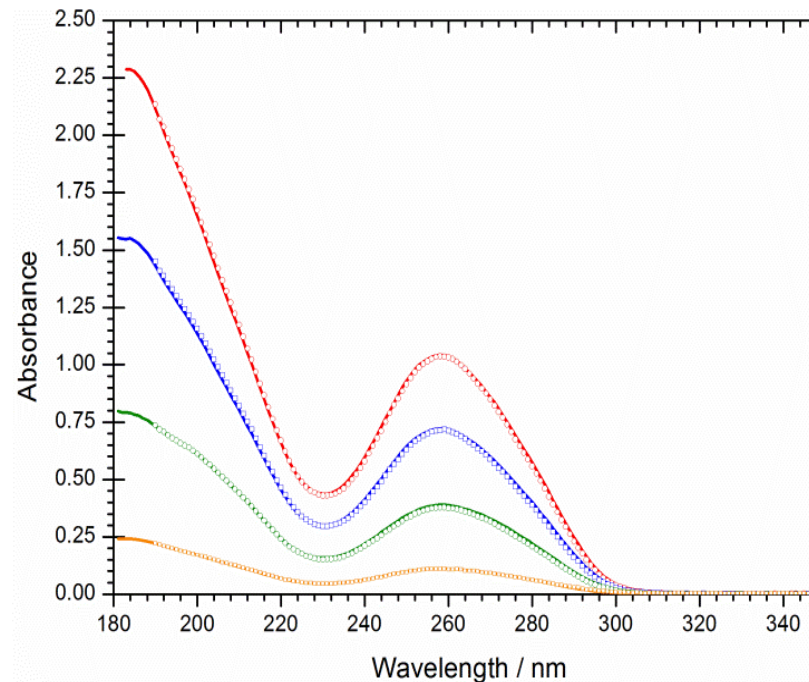
Only depends on a **single constant  $a$**  !  
Calibrate this using samples with know concentrations

# What does the HT tell us?

*Highly accurate simultaneous CD and Absorption measurements*

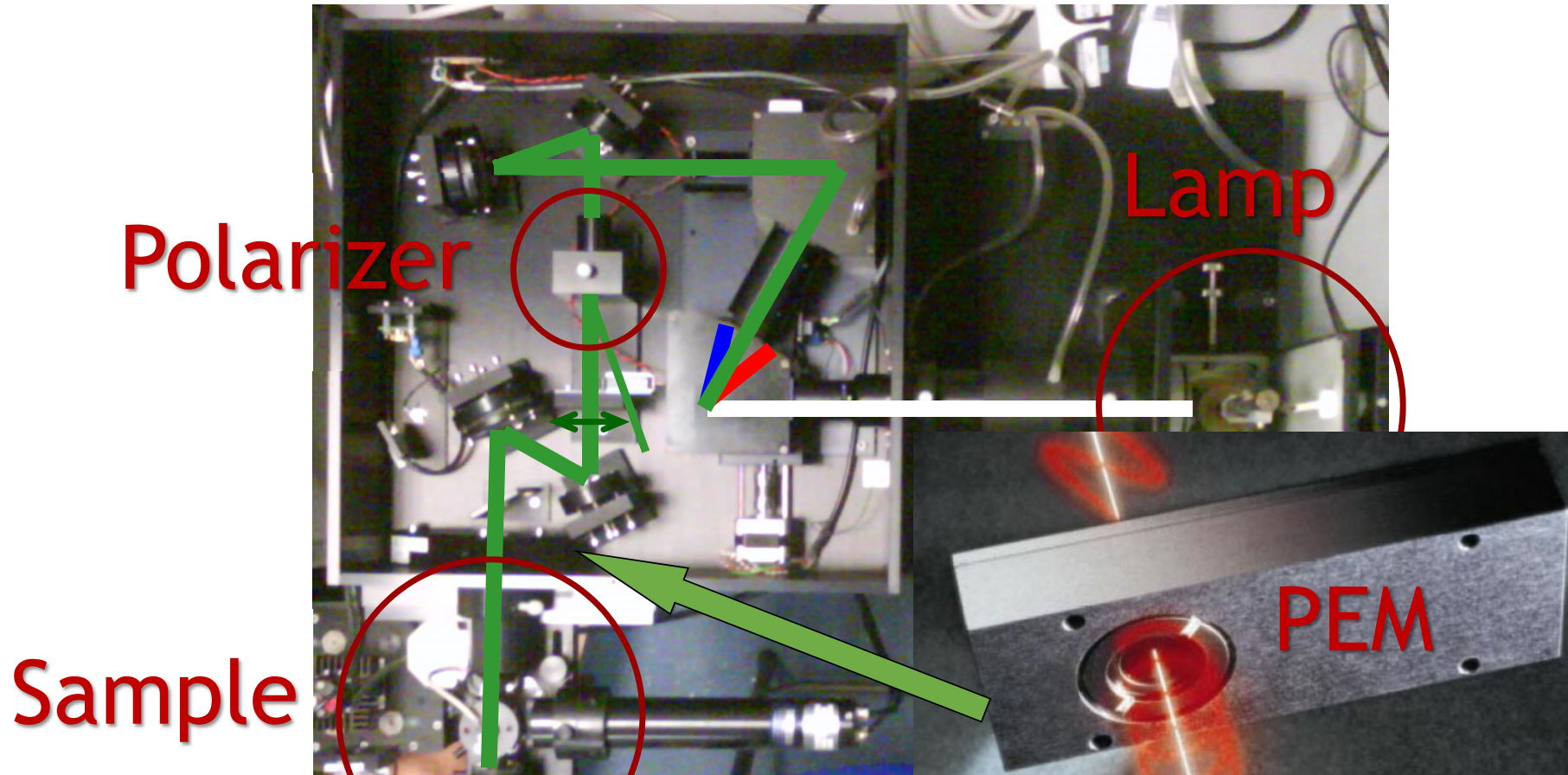
$$A(\lambda) = a \times \log \left( \frac{HT_{sample}(\lambda)}{HT_{baseline}(\lambda)} \right)$$

Full range of absorptions 0-2  
calibrated with a **single**  
detector parameter



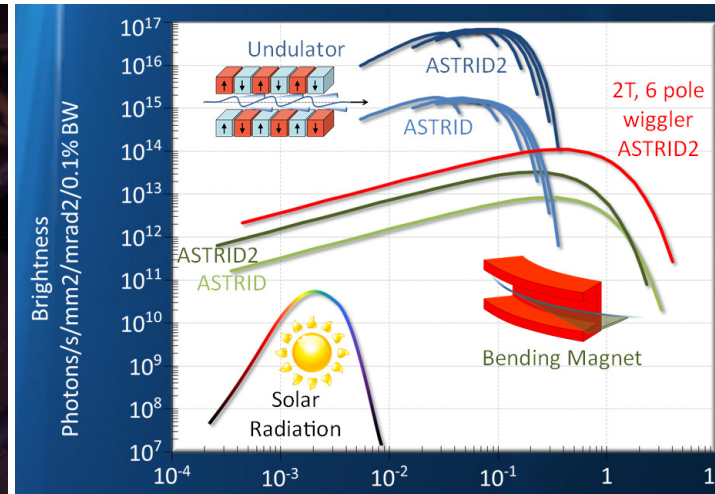
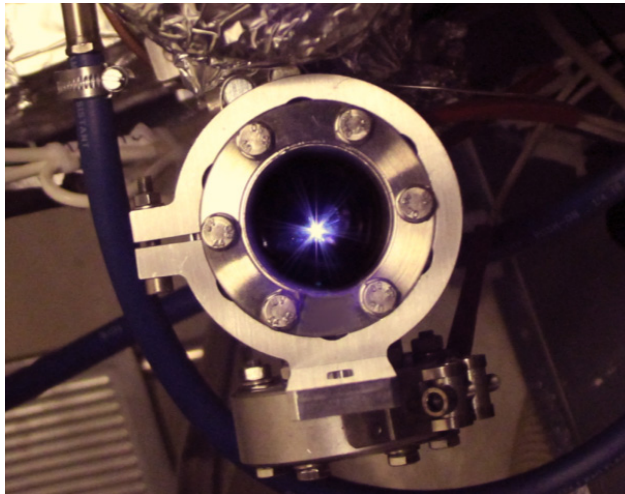
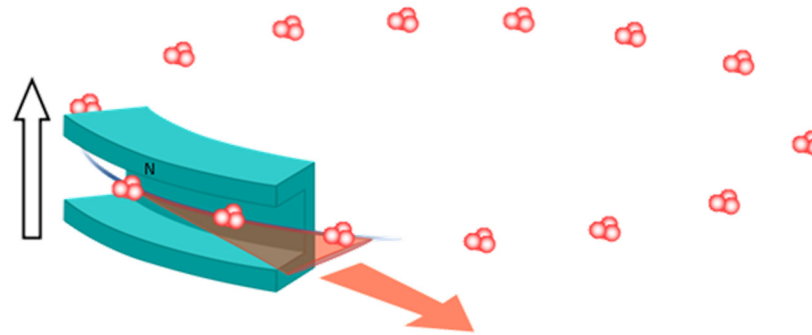
# Instrumentation

Conventional CD instrument (lamp based)

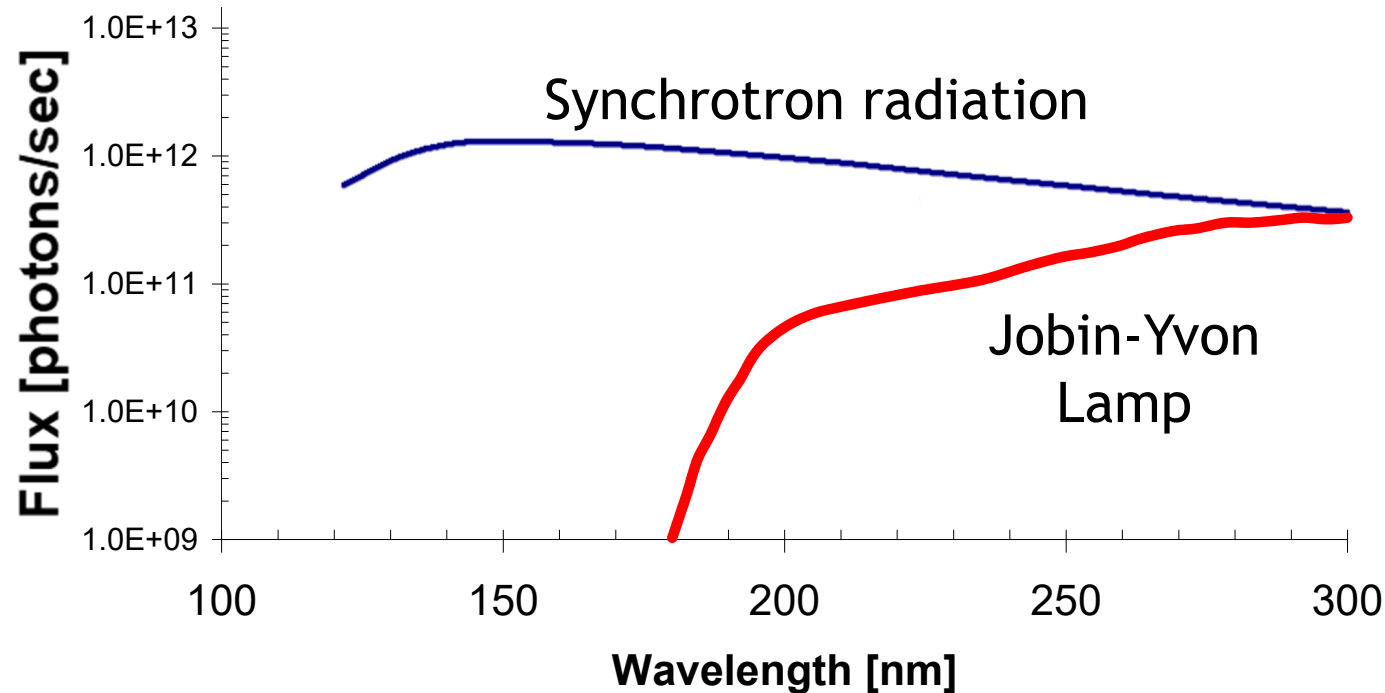


# Other options for light sources: Synchrotron Radiation (SR)

- UV light is well suited to examine molecules like proteins and DNA.
- Synchrotron radiation (SR) is emitted when charged particles are accelerated: We use electrons at relativistic speeds.
- The light is VERY intense.



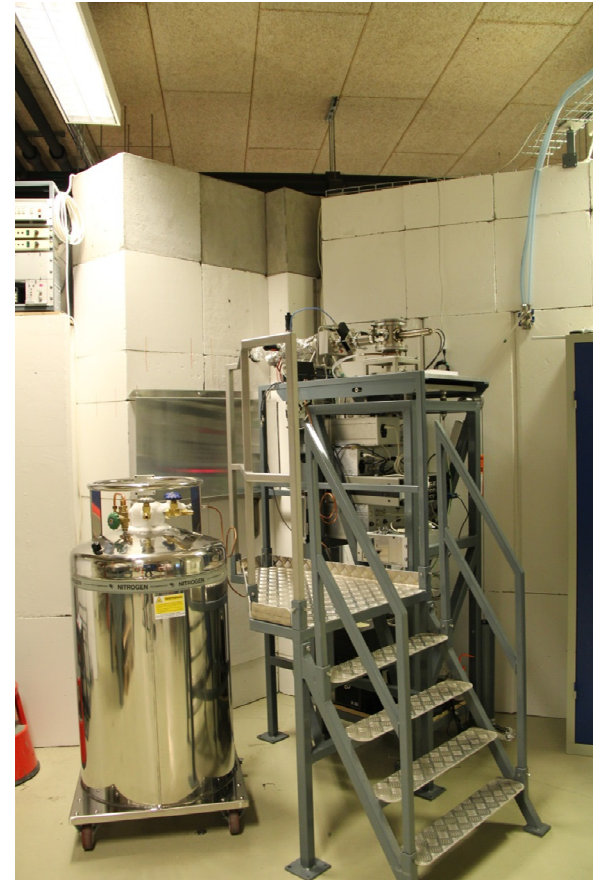
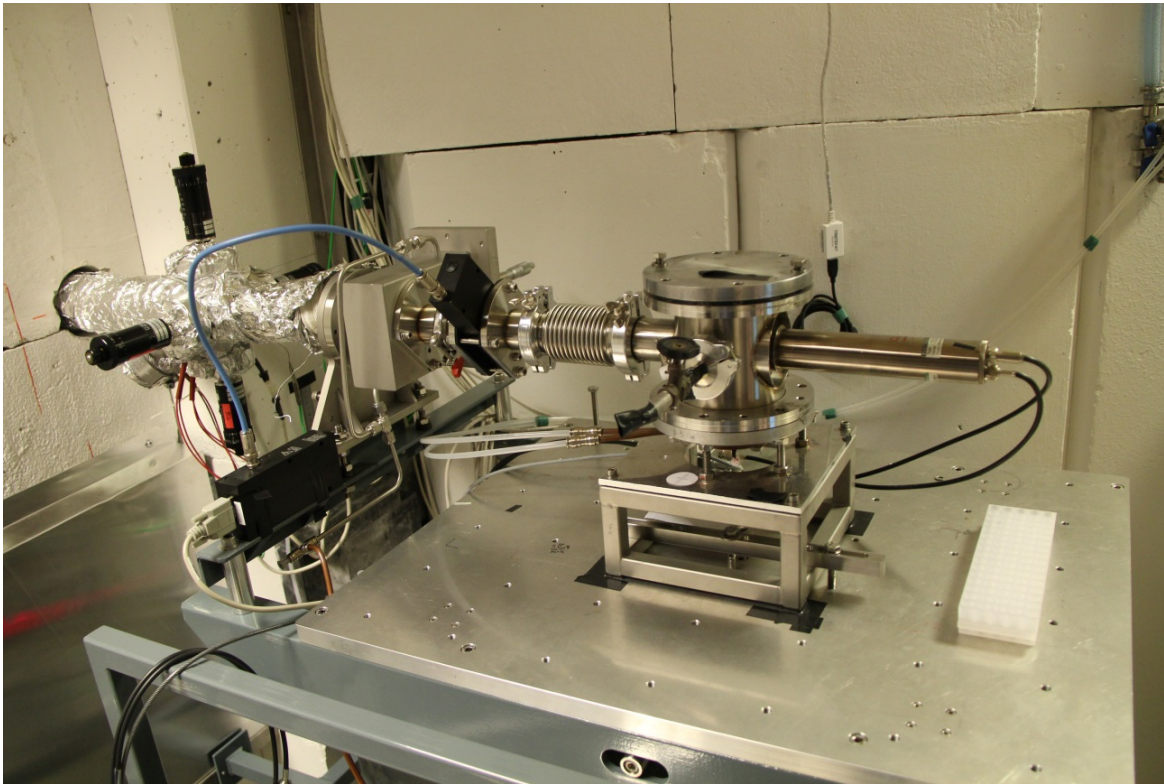
## Other options for light sources: SR vs lamps





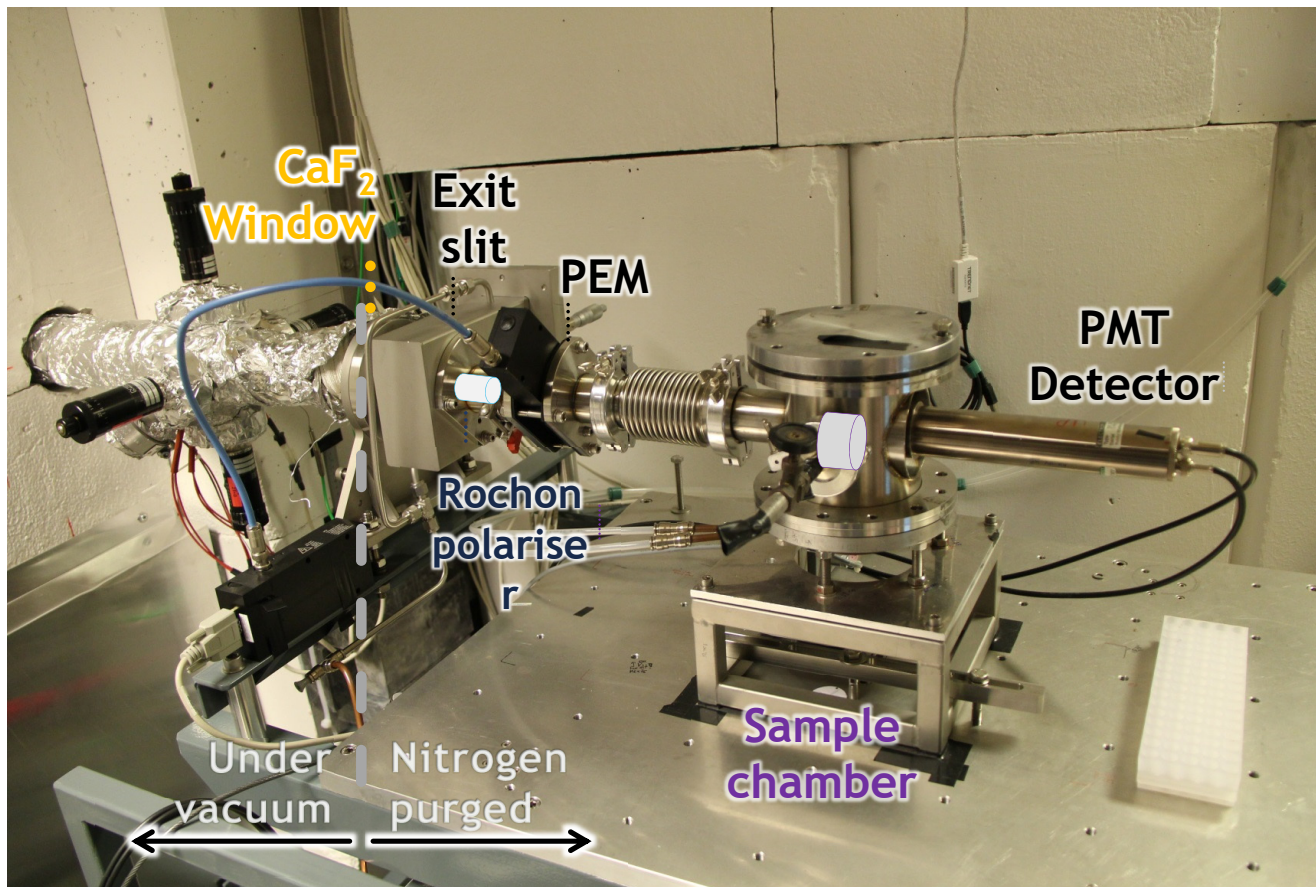
# AU-CD beam line on ASTRID2

## *Small and compact set-up*



This project has received funding from the European Union's Horizon 2020 research and innovation programme under grant agreement No 101004806

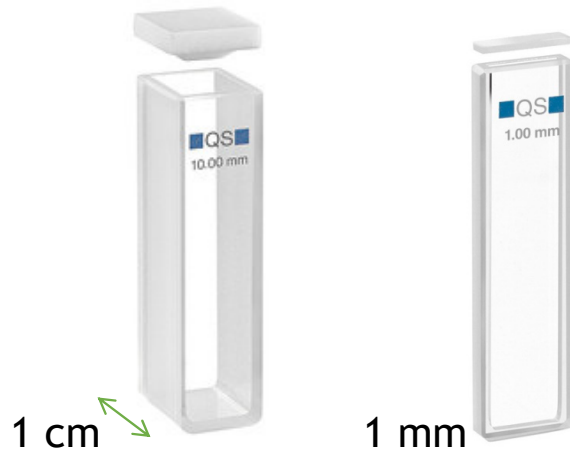
# AU-CD beam line on ASTRID2



# Measurement options...

## Sample holders - solvents

Cells made from Quartz suprasil with a cut-off of ~160 nm



### ► Closed cells

- Pathlengths 0.1 mm to several cm
- Temperature melts



### ► Open cells

- Shorter pathlengths (~0.01 mm or lower!)

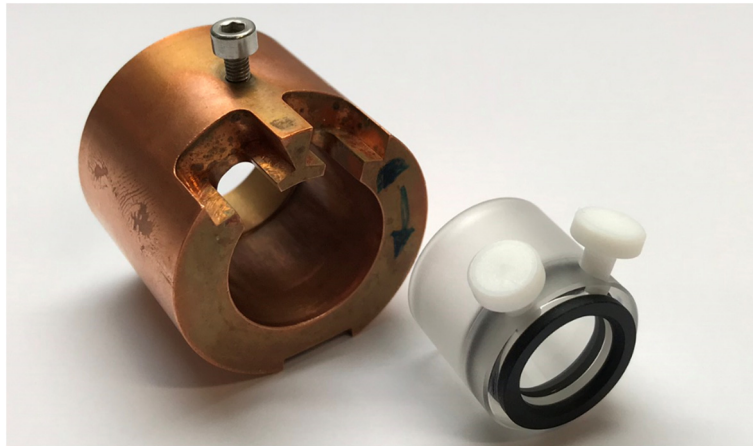


# Measurement options...

## Sample holders - solvents

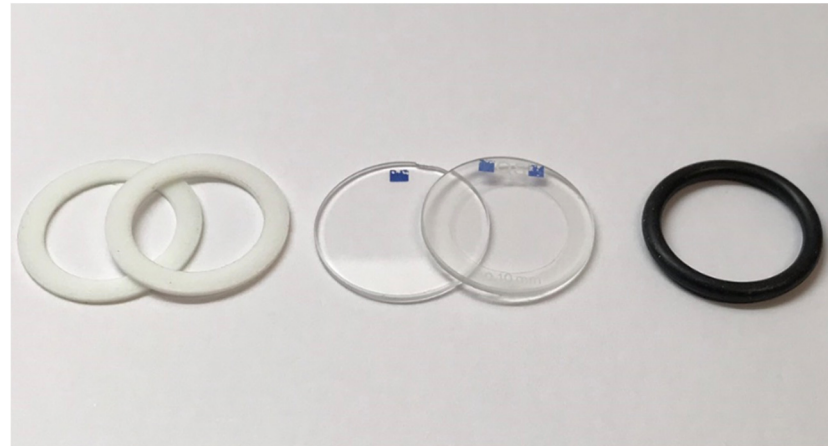
Cells made from Quartz suprasil with a cut-off of ~160 nm

We prefer to use round cells as they often have less stress and thus birefringence



### ► Closed cells

- Pathlengths 0.1 mm to several cm
- Temperature melts



### ► Open cells

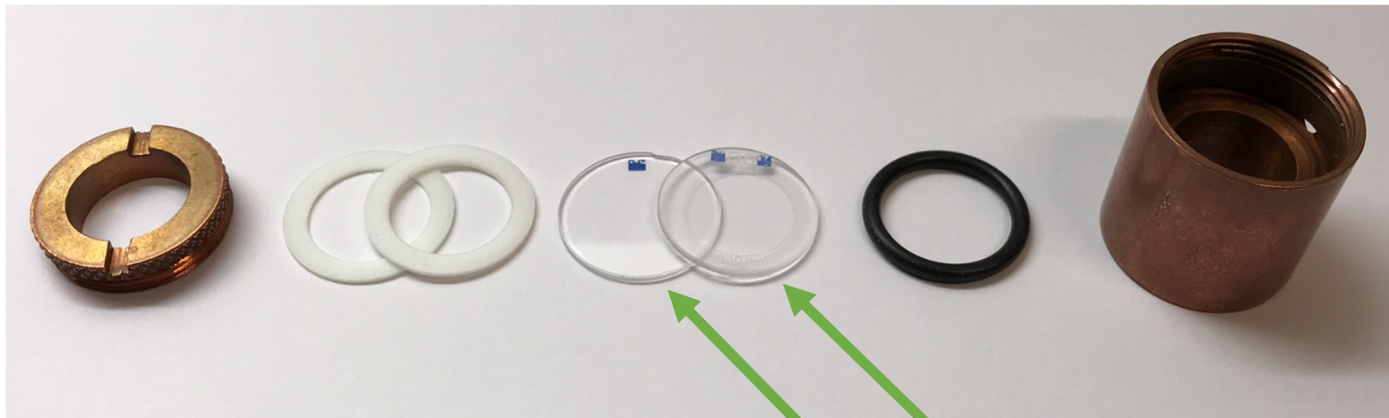
- Shorter pathlengths (~0.01 mm or lower!)
- $\text{CaF}_2$  cells, lower wavelength cut-off

# Measurement options...

## Sample holders - solvents

Cells made from Quartz suprasil with a cut-off of ~160 nm

We prefer to use round cells as they often have less stress and thus birefringence



### ► Open cells

- Shorter pathlengths (~0.01 mm or lower!)
- $\text{CaF}_2$  cells, lower wavelength cut-off



Flat plate

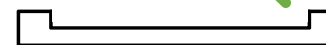


Plate with well

# A note on cell cleaning

Some cells are easier to clean than others



Large open cuvette:

Not so difficult to clean

Use:

Water



Open cells:

Easy to clean

Use:

Water and tissue

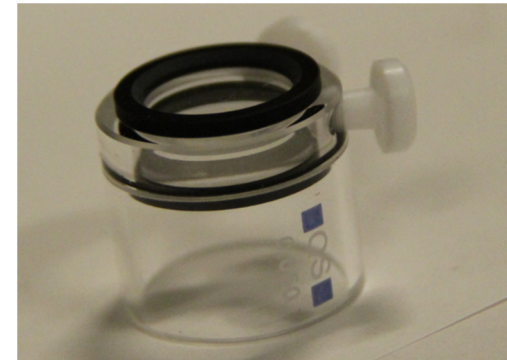


Closed cells:

Difficult to clean

Use:

Hellmanex 2% or 10% solution and heat

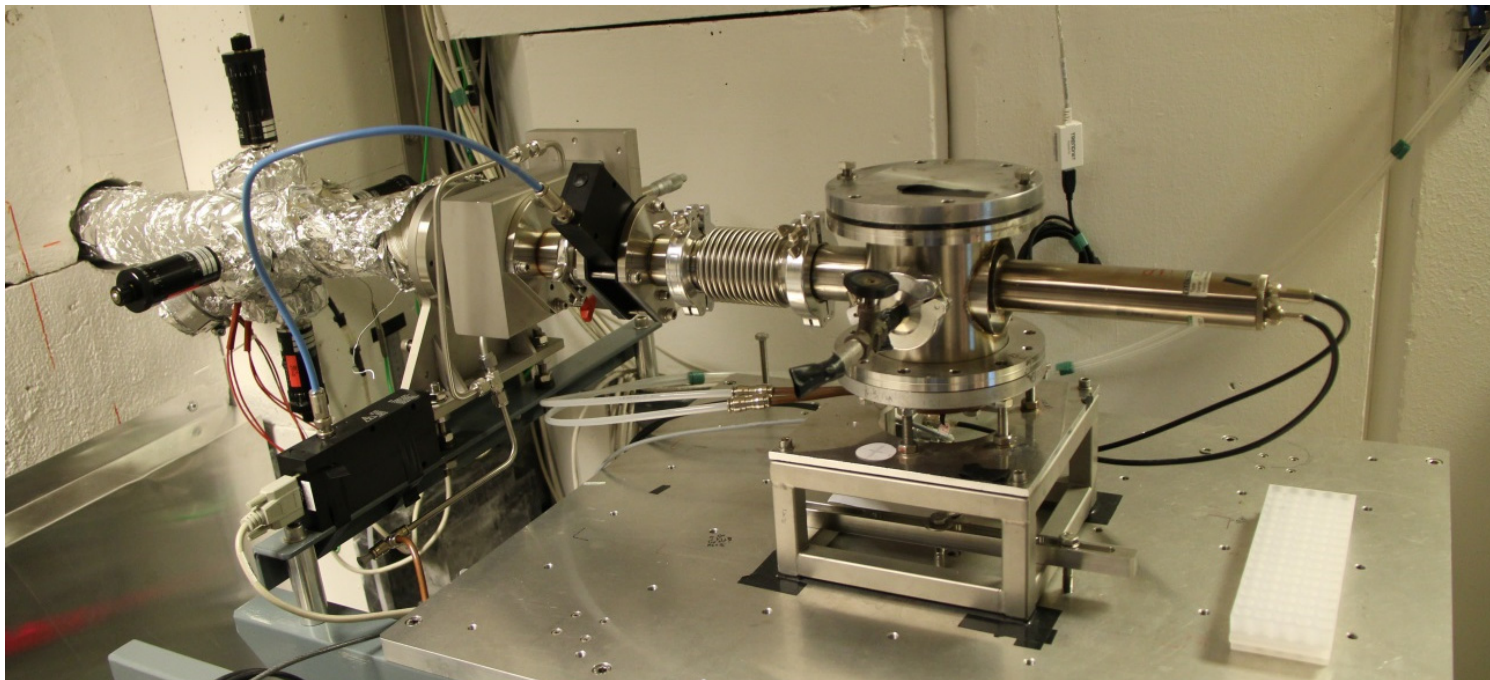


*Much more about this in the first Hands-on session*



# Measurement options...

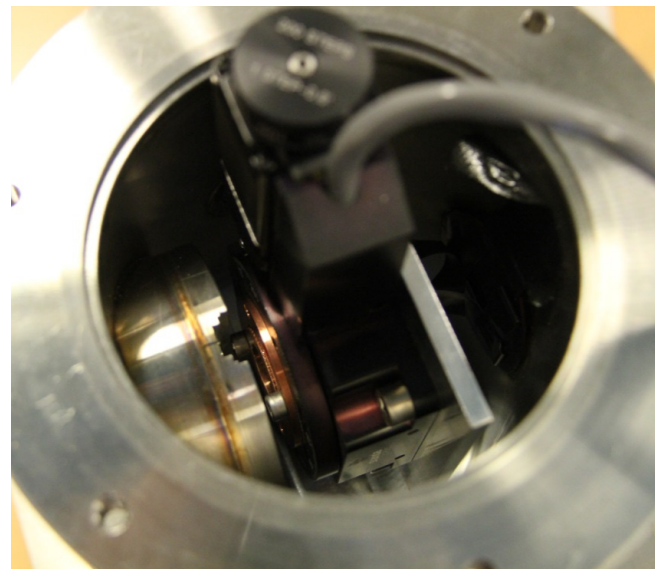
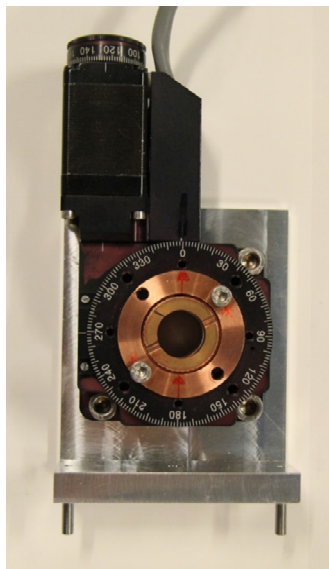
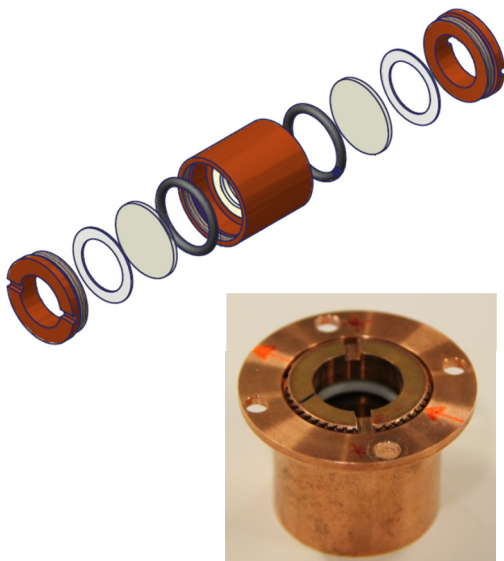
- Temperature scans 5 to 90C - fully automated and integrated into the scanning programme using a macro file.





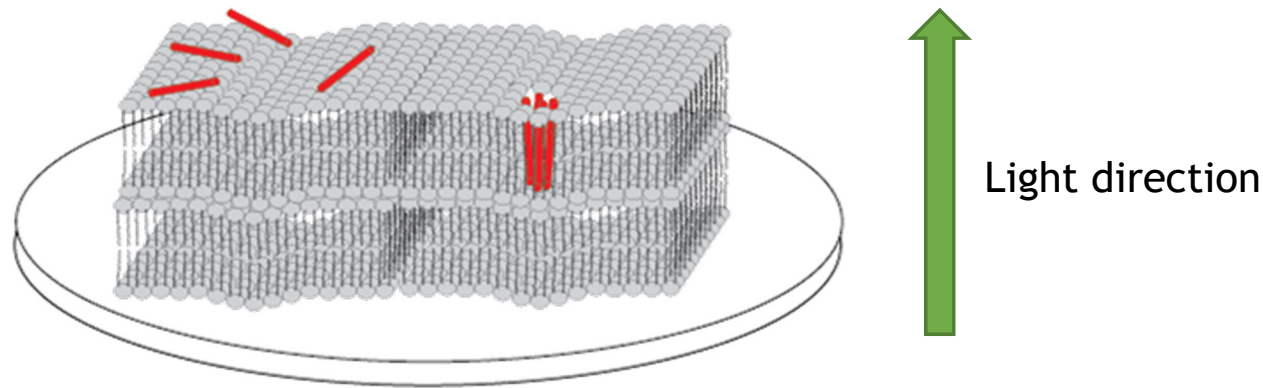
# Measurement options...

- ▶ Temperature scans 5 to 90C - fully automated and integrated into the scanning programme using a macro file.
- ▶ Rotational stage



# Why a rotation sample holder?

- ▶ Often used when studying films (solids) of samples
- ▶ A good example is insertion of peptides into lipid bilayers



Lipid bilayers (mixed with peptides) can be made on e.g. a quartz plate by drying a solution onto the plate.

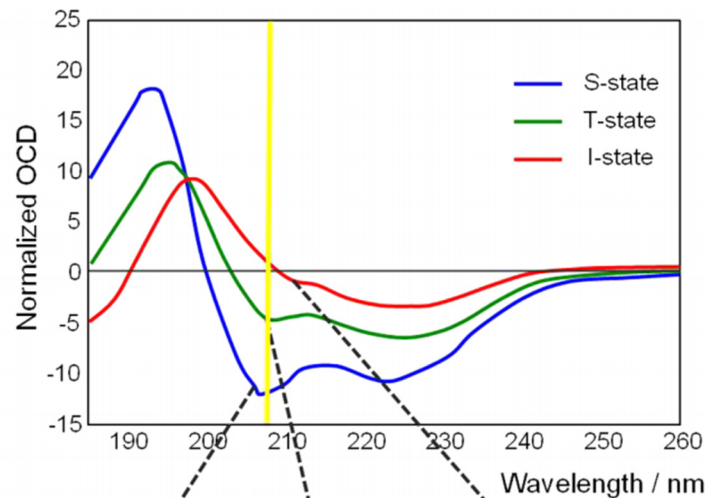
*Are the peptides inserted into the membrane or are they surface bound?*

This is a very different sample than a solution where the peptides are found in all directions with respect to the light

# Why a rotation sample holder?

## Oriented Circular Dichroism (OCD)

The CD spectrum of an helical peptide in the lipid bilayer depends on the orientation of the peptide.



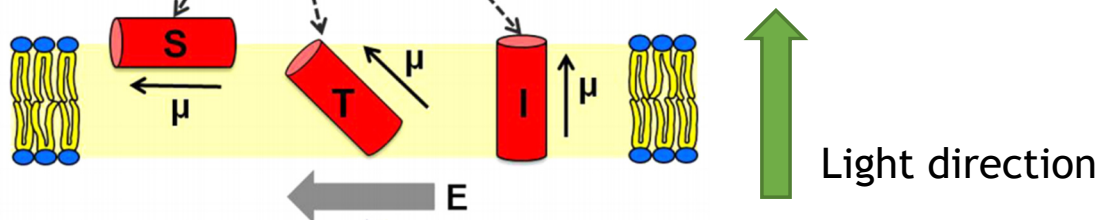
**S:** Surface bound peptide

**T:** Tilted peptide

**I:** Inserted peptide

The 208 nm transition ( $\mu$ ) is only active when parallel to the light electric field,  $E$  ( $\mu \parallel E$ )

$E$  is always perpendicular to the direction of the light



J. Bürck *et al.* *Acc. Chem. Res.* 2016, 49, 184–192

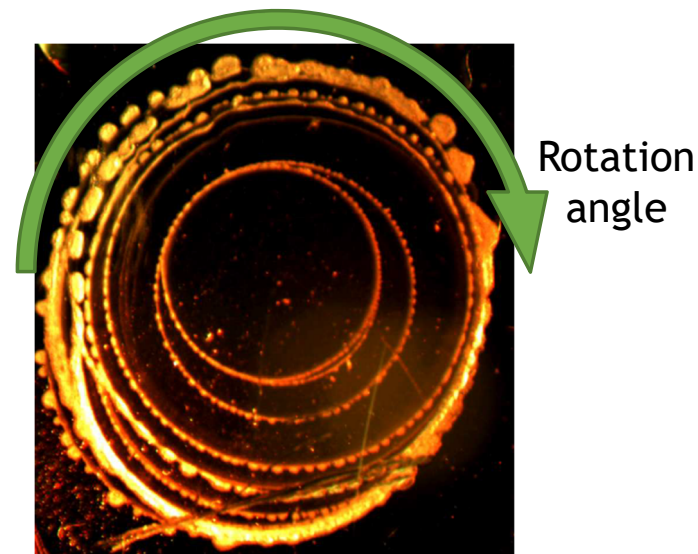
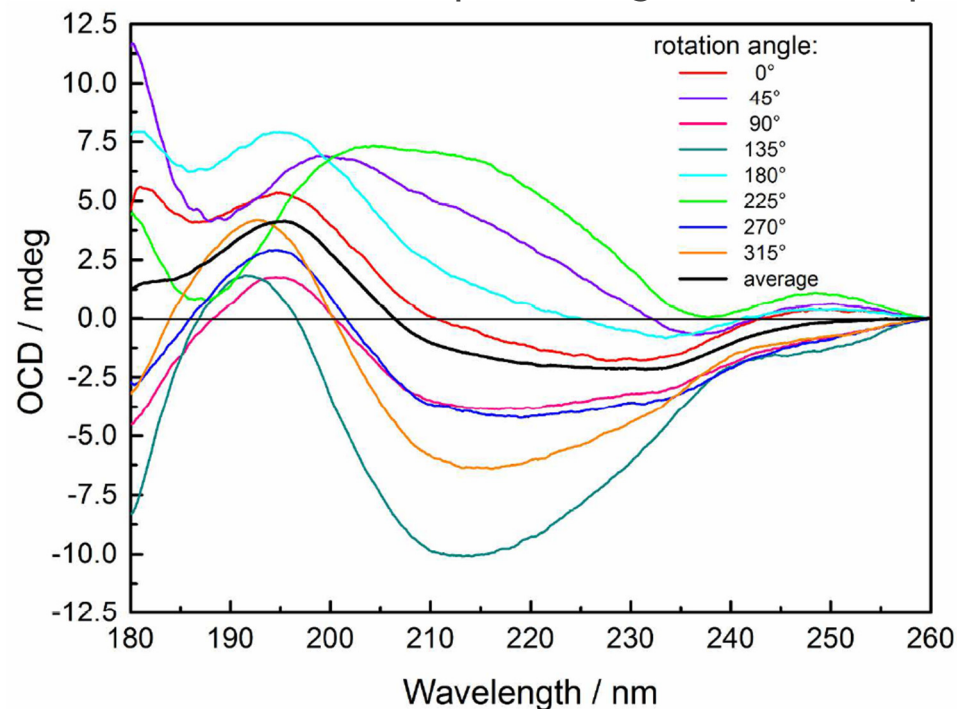


# Why a rotation sample holder?

## Oriented Circular Dichroism (OCD)

Such a (solid) sample is not necessarily very uniform, and may be a bit isotropic

- Shows up as changes in the CD spectrum with rotation angle



A not particularly uniform sample!

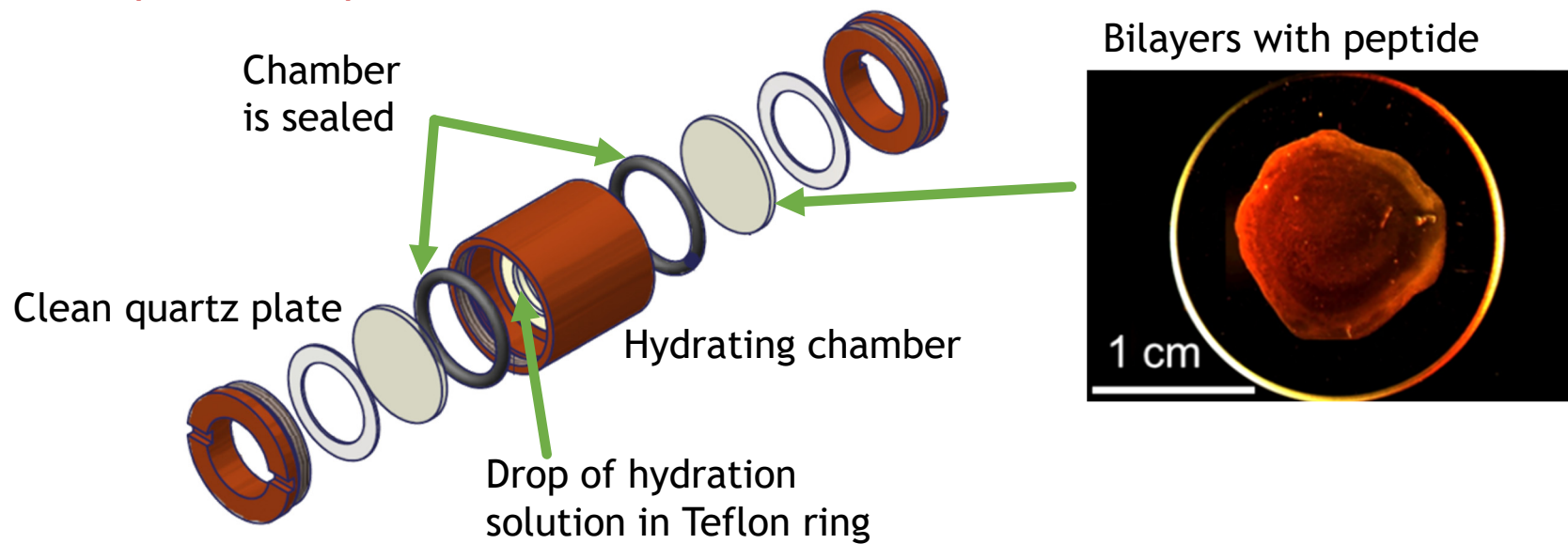
*The CD signal depends on sample angle, but averages to the correct spectrum!*

# Why a rotation sample holder?

## Oriented Circular Dichroism (OCD)

One important aspect of peptides in a bilayer is that they need to be kept humid and might also need to be hydrated after preparation for the peptides to insert into the membrane.

### Special sample holder

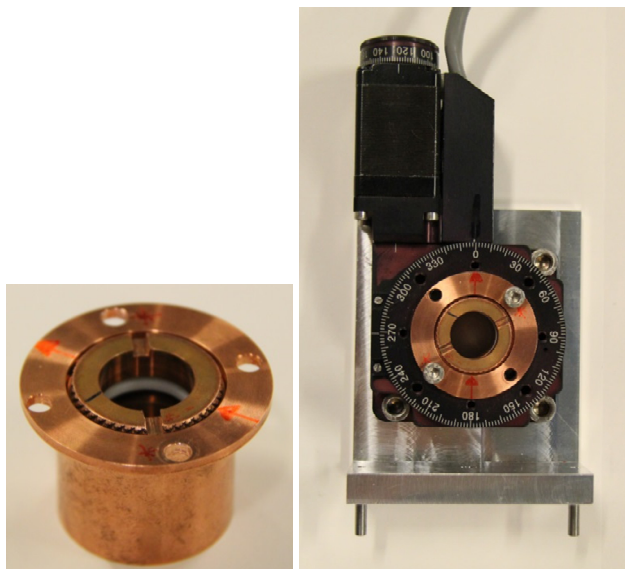


The humidity can be controlled with different salt solutions:  $\text{K}_2\text{SO}_4$  R.H. ~98 %  $\text{MgCl}_2$  R.H. ~ 33 %

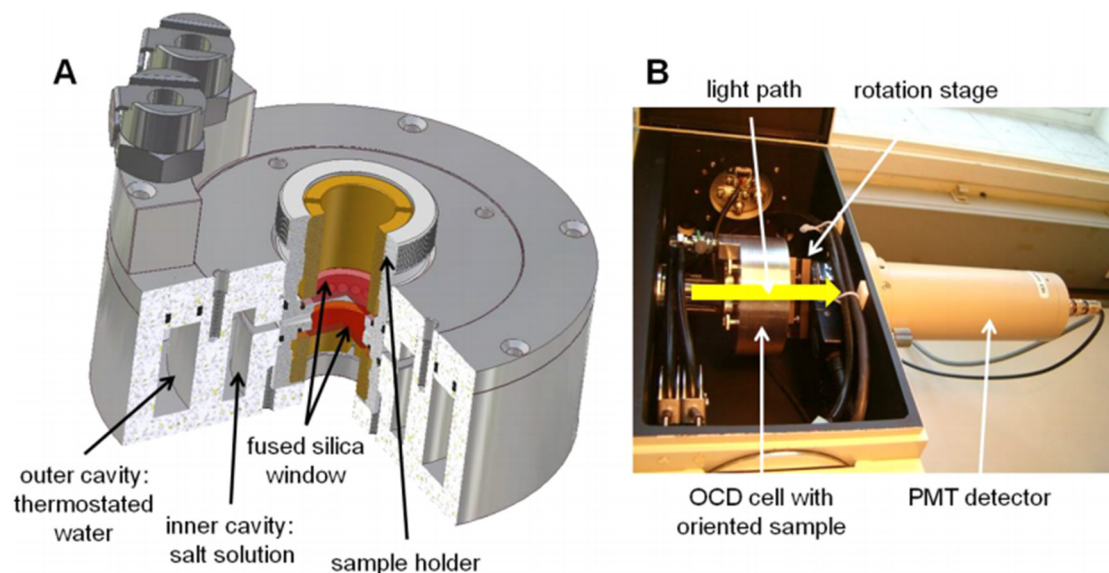
# Why a rotation sample holder?

## Oriented Circular Dichroism (OCD)

### Our setup for rotation



### KIT (DE) setup for a JASCO CD instrument

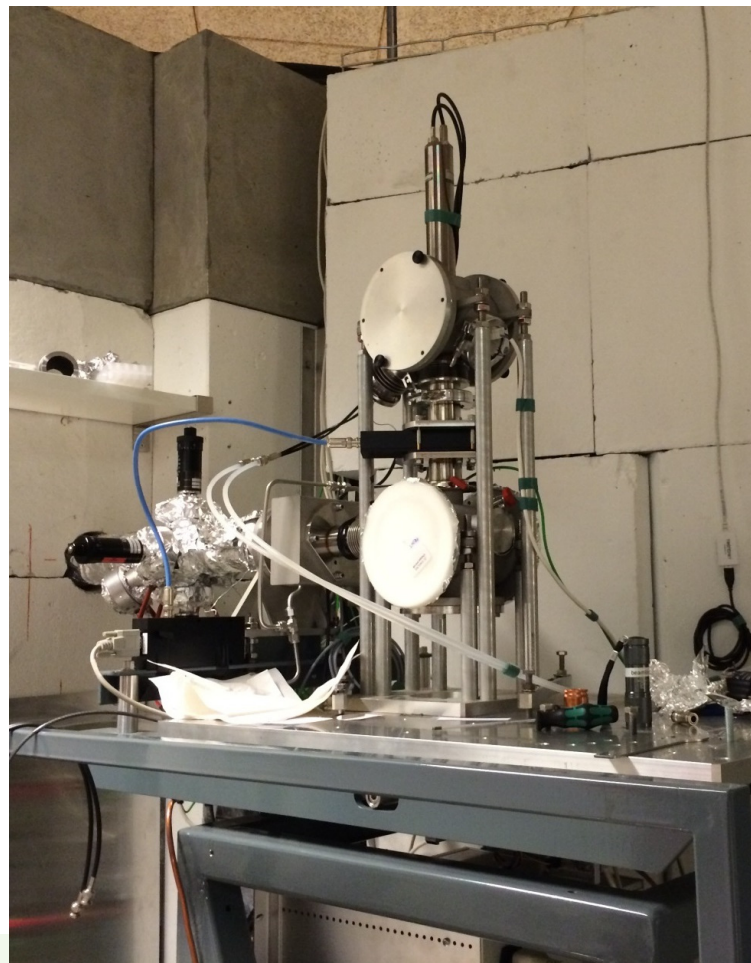
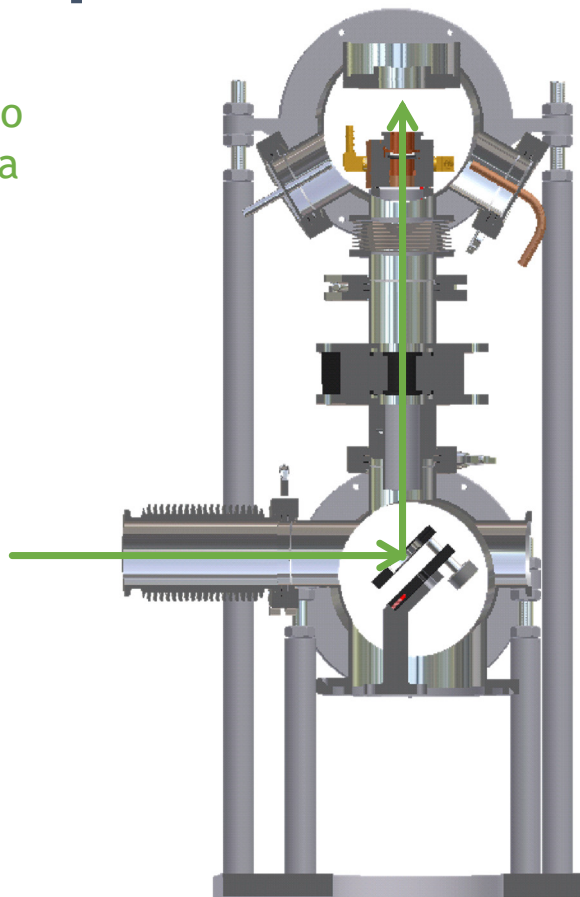


The hydrating sample holder is mounted in the rotational stage, and CD spectra are (automatically) acquired at many different angles.

## Measurement options...

# Periscope

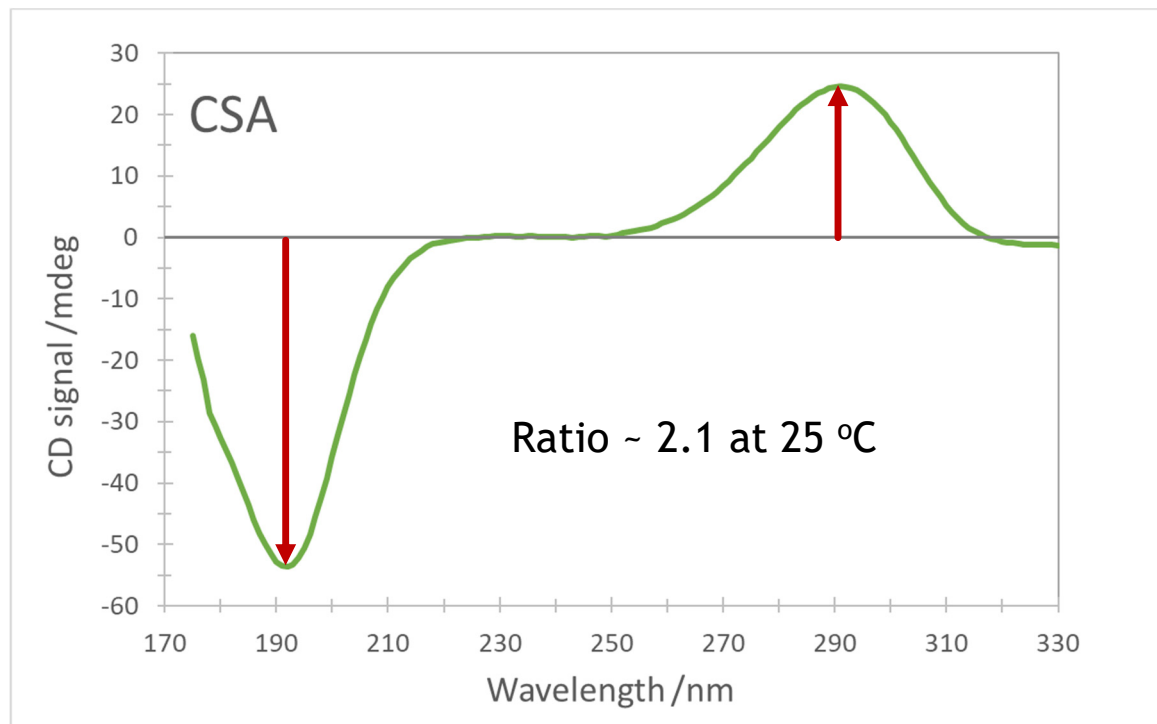
The periscope allows samples to be measured in a horizontal position





# Calibration and instrument health

As with any other instrument, it is vital that you keep your instrument well calibrated  
Check the calibration at the start of each day using (1S)-(+)-10-camphorsulfonic acid (CSA)



$$\Delta\epsilon_{290.5\text{nm}} = 2.36 \text{ M}^{-1} \text{ cm}^{-1}$$

$$\text{MW} = 232.30 \text{ g/mol}$$

*Use ~7 mg/ml in 0.1 mm cell*

## CSA concentration:

- CSA is hydroscopic
- Measure  $A_{285\text{nm}}$  in 1 cm cuvette
- Use  $\epsilon_{285} = 34.6 \text{ M}^{-1}\text{cm}^{-1}$
- Rule of thumb:

$$A_{285}(1\text{cm}) = 1 \text{ for } 6.71 \text{ mg/ml}$$

*Ratio must be above 2. Ratio is very temperature dependent: always use 25°C.*



# Calibration and instrument health

With time any spectrophotometer may develop stray light

- Stray light may pass through the sample
- If the absorption is high, the stray light hitting the detector may be significant
- The HT will become too low when keeping DC bias constant
  - Leads to CD features being measured too low!

The absorbance at the 192 nm CSA peak is relatively high

➤ Stray light shows up as the 192 / 290 ratio becoming low

*Daily measurements of the CSA peak values gives a track record of the health of the instrument:*

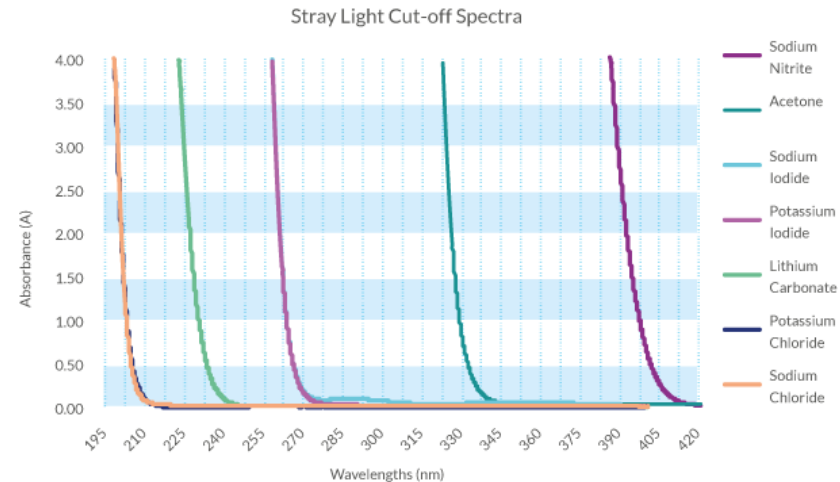
- The 192 / 290 ratio gives information about stray light changes over time
- The value of the 290 nm CSA CD signal gives information about the absolute calibration



# Calibration and instrument health

With time any spectrophotometer may develop stray light

E.g. Starna has reference salts for measuring stray light



Or you can make your own 12 g/l KCl solution:

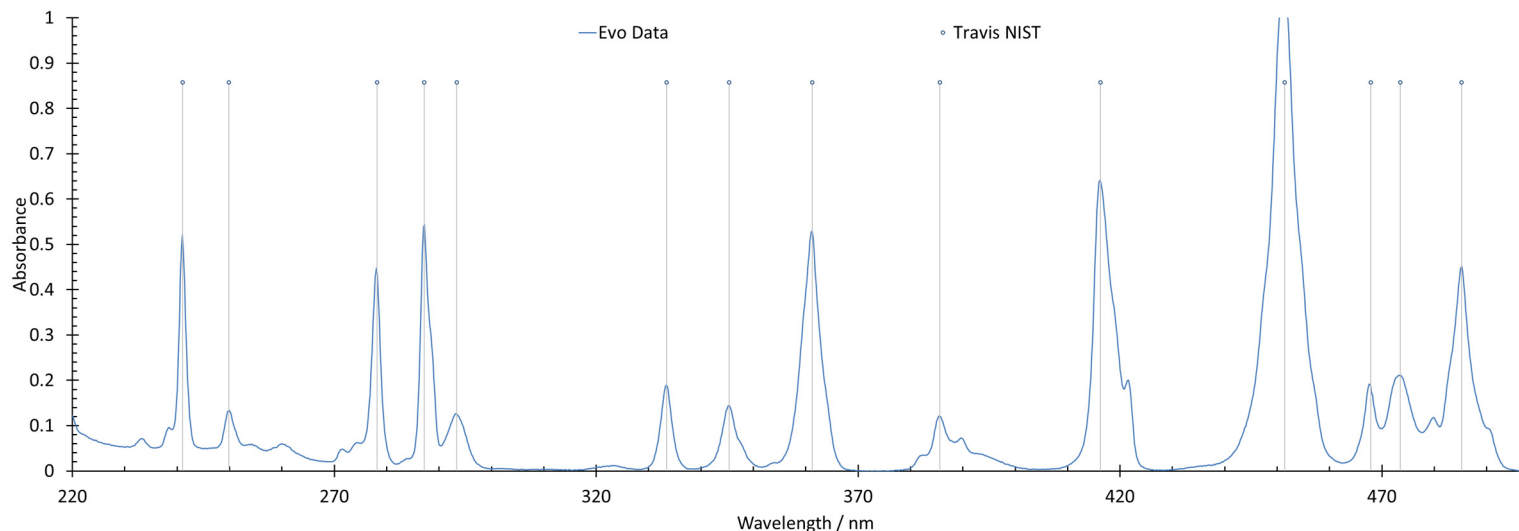
- Absorbance > 2.0 measured at 198 nm against a water reference



# Calibration and instrument health

## Check the wavelength calibration

- Daily by checking the position of the two CSA peaks (a rough estimate)
- Every few months using Holmium Oxide (HoOx)
  - 40 g/L solution of holmium oxide in 10% (v/v) perchloric acid
  - Store in sealed 10 mm cuvette



*Calibration data are found in:*

J.C. Travis *et al.* J. Phys. Chem. Ref. Data 34, 2005, 41-56